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# Radiation Diffusion in DRACO using Kershaw Difference Scheme 

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## 1. Introduction

Radiation hydrodynamics is the study of very high energy density plasma fluid systems where heat is transferred predominantly by the transfer of x-ray radiation. These systems are found in astrophysics and on earth in high energy density plasmas like those found in laser fusion targets [1]. The mathematical treatment of radiation hydrodynamics is extremely complex because it is a combination of fluid dynamics and radiative transfer. Each of these subjects themselves is very involved and in combination the problem offers no useful analytical mathematical solutions. Therefore one must immediately resort to numerical solutions to very nonlinear sets of coupled partial differential equations. The opacities used as coefficients in these equations are nonlinear functions of the plasma energy density (density and temperature), rendering the solutions even more difficult to compute. Because one is forced to simultaneously simulate the fluid motion under the influence of radiative heat transfer and the heat transfer in moving background plasma, computer resources limit the choice of radiative transfer model. For this reason, the diffusion approximation is often chosen over more accurate transport approximations to model the radiative transfer, even in circumstances where the diffusion approximation is not strictly valid. Therefore the diffusion approximation is modified with a "fix-up" called a flux limit.

The equations of radiation hydrodynamics can be written as [2]

$$
\begin{equation*}
\frac{\partial I}{c \partial t}-\Omega \cdot \nabla I=S(v)-\sigma(v) I+\int_{0}^{\infty} d v^{\prime} \int_{4 \pi} d \Omega^{\prime} \frac{v}{v^{\prime}} \sigma_{s}\left(v^{\prime}-v, \Omega^{\prime} \cdot \Omega\right) I\left(v^{\prime}, \Omega^{\prime}\right) \tag{1.1}
\end{equation*}
$$

where we have omitted the fluid equations for the conservation of mass and momentum and only included the conservation of energy equation, because this is the fluid equation that couples directly to the radiative transfer equation. In this report we will focus only on the solution of the radiative transfer equation in the diffusion approximation.

The radiative transfer equation is treated in the multi-frequency approximation where the frequency dependence of the specific intensity is treated in a group structure with average opacities computed for each group. Each of these group equations is uncoupled from the others. In fact they are coupled through the emission term that appears in both the radiative transfer equations and the plasma energy density equation. This coupling is broken using an explicit treatment of the solution of these equations, that is the radiative transfer equations are updated from $t^{n}$ to $t^{n+1}$ using values of plasma parameters at $\mathrm{t}^{\mathrm{n}}$. The plasma energy is updated using an emission term evaluated at $\mathrm{t}^{\mathrm{n}}$ and an absorption term evaluated at $\mathrm{t}^{\mathrm{n}+1}$. This time centering inconsistency allows sequential solution of these equations, but introduces errors and possibly numerical instabilities that must be controlled.

Taking moments of the radiative transfer equations with respect to their angular dependence and truncating after zeroth and first moments produces k equations and $\mathrm{k}+1$ unknown functions. This can also be viewed as expanding the specific intensity in a power series in the angular variable and retaining only the zeroth and first order terms. In either case, assuming a transport law that relates the gradient of the radiation energy density to the radiation flux in the form of Fick's Law closes the resulting radiation energy density equation:

$$
\begin{equation*}
\vec{F}=-D \nabla I, \tag{1.2}
\end{equation*}
$$

where D is the flux-limited diffusion coefficient. The diffusion form of the radiative transfer equation is [2]

$$
\begin{equation*}
\frac{\partial I_{g}}{c \partial t}-\nabla \cdot D \nabla I_{g}=4 \pi \sigma_{a}^{\prime} B_{v}-\sigma_{a}^{\prime} I_{g} \tag{1.3}
\end{equation*}
$$

It is this equation, in multi-frequency form, that is solved in DRACO. The emission and the absorption terms that couple the radiation diffusion equations to the plasma energy density
equation are simply cell-wise quantities. There is no spatial finite differencing process involved in representing these quantities.

The major challenge in solving these equations is the spatial finite difference approximation to the diffusion term. This is the term that transports the radiation from one cell to another. Furthermore, DRACO [3] is a 2D lagrangian code, so the finite difference mesh is composed of quadrilateral cells in a structured mesh. By this we mean that cells map to a regular logical index mesh, but the geometric relationship of a cell and its surrounding cells can be arbitrarily distorted as shown in Figure 1. DRACO also uses either x-y or r-z coordinates. For x-y coordinates and orthogonal cell boundaries there are numerous finite difference approximations to the so-called diffusion operator. For r-z coordinates and arbitrary quadrilateral cells, there are few published finite difference schemes.


Fig. 1.1 Logical coordinates and quadrant numbering for a quadrilateral cell on a logically rectangular mesh.

David Kershaw developed the finite difference scheme that is best documented in the published literature and we hereafter refer to this as the "Kershaw scheme" [4].

In Section 2, we give a description of the Kershaw scheme by following his paper. We also discuss the boundary condition treatment and the different types of the flux limiters. Detailed information about the matrix and the corresponding symbols appearing in the diffusion equation is also given. In Section 3, we present the results of simple test problems in the contour graphs. We
also compare the numerical results to the exact solutions for several problems that have analytical solutions. In Section 4, we list some variables related to the code implementation and the input options for users.

## 2. Kershaw difference scheme for the diffusion equation on an arbitrary r-z quadrilateral grid

### 2.1 Differencing the diffusion operator [4]

To difference the diffusion operator, one observes from Fig. 1.1 that

$$
\begin{equation*}
f=f(R(K, L), Z(K, L)) \tag{2.1}
\end{equation*}
$$

$R$ and $Z$ are the familiar Eulerian coordinates, $R$ being either Cartesian or cylindrical and $Z$ always being Cartesian. (The following derivations will be carried out in cylindrical coordinates.) $K$ and $L$ are the mesh point indices coordinates, with $K=1, \ldots, K M A X$ and $L=1, \ldots, L M A X$.

The expression $\nabla \cdot D \nabla f$ in cylindrical coordinates is:

$$
\begin{equation*}
\frac{1}{R} \frac{d}{d R} D R \frac{d f}{d R}+\frac{d}{d Z} D \frac{d f}{d Z} \tag{2.2}
\end{equation*}
$$

If $K$ and $L$ are treated as continuous variables of $R$ and $Z$, one can transform the coordinates from $(R, Z)$ to ( $K, L$ ) using the relationships

$$
\begin{align*}
& \frac{d f}{d R}=\frac{d f}{d K} \frac{d K}{d R}+\frac{d f}{d L} \frac{d L}{d R}  \tag{2.3}\\
& \frac{d f}{d Z}=\frac{d f}{d K} \frac{d K}{d Z}+\frac{d f}{d L} \frac{d L}{d Z}
\end{align*}
$$

We can define the area Jacobian of the cell as

$$
\begin{equation*}
j=\frac{d(R, Z)}{d(K, L)}=\frac{d R}{d K} \cdot \frac{d Z}{d L}-\frac{d Z}{d K} \cdot \frac{d R}{d L} . \tag{2.4}
\end{equation*}
$$

We define

$$
\begin{equation*}
R_{K}=\binom{\frac{d R}{d K}}{\frac{d Z}{d K}} \quad R_{L}=\binom{\frac{d R}{d L}}{\frac{d Z}{d L}} \tag{2.5}
\end{equation*}
$$

Then working through the partial derivatives, one gets

$$
\begin{align*}
\nabla \cdot(D \nabla f) & =\frac{1}{R j}\left\{\frac{d}{d K}\left[\frac{D R\left(R_{L}\right)^{2}}{j} \frac{d f}{d K}\right]+\frac{d}{d L}\left[\frac{D R\left(R_{K}\right)^{2}}{j} \frac{d f}{d L}\right]\right\} \\
& -\frac{1}{R j}\left\{\frac{d}{d K}\left[\frac{D R\left(R_{K} \cdot R_{L}\right)}{j} \frac{d f}{d L}\right]+\frac{d}{d L}\left[\frac{D R\left(R_{K} \cdot R_{L}\right)}{j} \frac{d f}{d K}\right]\right\} \tag{2.6}
\end{align*}
$$

The difference approximation to this equation is derived using a technique based on a variational formulation of the solution. Using the identity

$$
\nabla \cdot(f D \nabla f)=f \nabla \cdot D \nabla f=D(\nabla f)^{2}
$$

one can write the equation

$$
\begin{align*}
& \int R d R d Z f \nabla \cdot(D \nabla f)=\int R d R d Z \nabla \cdot(f D \nabla f)-\int R d R d Z D(\nabla f)^{2}  \tag{2.7}\\
& \int R d R d Z f \nabla \cdot(D \nabla f)=\int R d R d Z \nabla \cdot(f D \nabla f)-\int R d R d Z D(\nabla f)^{2} \\
& \int R d R d Z \nabla \cdot(f D \nabla f)=0,
\end{align*}
$$

since by the divergence theorem

$$
\int R d R d Z \nabla \cdot(f D \nabla f)=\int f D d S \cdot \nabla f
$$

and on the boundary of the problem, it is assumed that $f=0$ or $d S \cdot \nabla f=0$. Occasionally one has $f=f_{B} \neq 0$ on the boundary. When this occurs, the terms in $(A f)_{K, L}$ that involve $f_{B}$ are now known and can be moved to the right-hand side of Eq. 2.5 when it is finite differenced. Then the remaining matrix $A$, which operates on the vector of unknowns, is the same as if $f=0$ on the boundary.

Then Eq. (2.7) is written

$$
\begin{equation*}
\int R d R d Z f \nabla \cdot(D \nabla f)=-\int R d R d Z D(\nabla f)^{2} . \tag{2.8}
\end{equation*}
$$

The finite difference analogue of the differentials in Eq. (2.8) is

$$
\begin{equation*}
\sum_{K, L} f_{K, L}(A f)_{K, L} V_{K, L}=-\sum_{K, L}|B f|_{K, L}^{2} \tag{2.9}
\end{equation*}
$$

where

$$
V_{K, L}=R j_{K, L}=\text { zone volume } / 2 \pi .
$$

and $A$ is the matrix to be determined. The matrix $B$ will be used to define $A$.

### 2.2 Definition of the elements of the finite difference operator

The form of $(B f)$ is chosen so that:

$$
\begin{equation*}
-\int R d R d Z D(\nabla f)^{2} \approx-\frac{1}{4} \sum_{K, L} \sum_{i=1}^{4}\left|B^{i} f\right|_{K, L}^{2} \tag{2.10}
\end{equation*}
$$

The reason for choosing to divide by $1 / 4$ is not immediately obvious, but will become apparent later on. The left hand side integral is transformed to ( $K, L$ ) coordinates (again treating $K$ and $L$ as continuous variables) to get

$$
\begin{equation*}
-\int R d R d Z D(\nabla f)^{2}=-\int d K d L\left[\left(\frac{D R}{j}\right)^{\frac{1}{2}}\left(R_{L} \frac{d}{d K}-R_{K} \frac{d}{d L}\right) f\right]^{2} \tag{2.11}
\end{equation*}
$$

From this, one can define:

$$
\begin{equation*}
(B f)_{K, L} \approx\left(\frac{D R}{j}\right)^{1 / 2}\left(R_{L} \frac{d f}{d K}-R_{K} \frac{d f}{d L}\right) . \tag{2.12}
\end{equation*}
$$

This can be written:

$$
(B f)_{K, L} \simeq k_{K, L} R_{L}-\ell_{K, L} R_{K}
$$

where

$$
\begin{align*}
& k_{K, L} \simeq\left(\frac{D R}{j}\right)^{\frac{1}{2}} \frac{d f}{d K} \\
& \ell_{K, L} \simeq\left(\frac{D R}{j}\right)^{\frac{1}{2}} \frac{d f}{d L} \tag{2.13}
\end{align*}
$$

k and $\ell$ are face centered quantities because they involve $\frac{d f}{d K}$ and $\frac{d f}{d L}$, so they can be differenced

$$
\begin{align*}
k_{K, L} & =\Sigma_{K, L}\left(f_{K+1, L}-f_{K, L)}\right)  \tag{2.14}\\
\ell_{K, L} & =\Lambda_{K, L}\left(f_{K, L+1}-f_{K, L}\right)
\end{align*}
$$

where $\Sigma_{K, L}=(D R / j)^{1 / 2}$ and is suitably averaged between the zones $(K, L)$ and $(K+1, L)$, and $\Lambda_{K, L}=(D R / j)^{1 / 2}$ is suitably averaged between the zones $(K, L)$ and $(K, L+1)$. Since $K$ and $L$ are actually discrete variables, the partial derivatives have been differenced as

$$
\begin{align*}
& \frac{d f}{d K}=\frac{f_{K+1, L}-f_{K, L}}{(K+1)-K}=f_{K+1, L}-f_{K, L} \\
& \frac{d f}{d L}=\frac{f_{K+L+1}-f_{K, L}}{(L+1)-L}=f_{K, L+1}-f_{K, L} \tag{2.15}
\end{align*}
$$

One choice for a suitable $\Sigma_{K, L}$ and $\Lambda_{K, L}$ is given by

$$
\begin{equation*}
\Sigma_{K, L}^{2}=\frac{\left(R_{K, L}+R_{K, L-1}\right)}{\frac{j_{K, L}}{D_{K, L}}+\frac{j_{K+1, L}}{D_{K+1, L}}} \quad \Lambda_{K, L}^{2}=\frac{\left(R_{K, L}+R_{K, L+1}\right)}{\frac{j_{K, L}}{D_{K, L}}+\frac{j_{K, L+1}}{D_{K, L+1}}} \tag{2.16}
\end{equation*}
$$

The terms $\mathrm{k}_{\mathrm{K}, \mathrm{L}}$ and $\ell_{\mathrm{K}, \mathrm{L}}$ are represented in diagram form in Figs. 2.1a,b.


Fig. 2.1a. $\mathrm{k}_{\mathrm{K}, \mathrm{L}}$


Fig. 2.2b. $\ell_{\mathrm{K}, \mathrm{L}}$

Remember that

$$
R_{K}=\binom{\frac{d R}{d K}}{\frac{d Z}{d K}} \quad R_{Z}=\binom{\frac{d R}{d L}}{\frac{d Z}{d L}}
$$

and let $R_{K, L}=\left(R_{K, L}, Z_{K, L}\right)$, then $\left(R_{K}\right)_{K, L}$ and $\left(R_{L}\right)_{K, L}$ are differenced

$$
\begin{equation*}
\left(R_{K}\right)_{K, L}=\frac{\left(R_{K, L}+R_{K, L-1}\right)}{2}-\frac{\left(R_{K-1, L}+R_{K-1, L-1}\right)}{2} \tag{2.17}
\end{equation*}
$$



$$
\begin{equation*}
\left(R_{L}\right)_{K, L}=\frac{\left(R_{K, L}+R_{K-1, L}\right)}{2}-\frac{\left(R_{K, L-1}+R_{K-1, L-1}\right)}{2} \tag{2.18}
\end{equation*}
$$



Then $\left(B^{l} f\right)$ is chosen to be

$$
\begin{gather*}
\left(B^{1} f\right)_{K, L}=k_{K, L}\left(R_{L}\right)_{K, L}-\ell_{K, L}\left(R_{K}\right)_{K, L}  \tag{2.19}\\
\left(B^{1} f\right)_{K, L}=\sum_{K, L}\left(f_{K+1, L}-f_{K, L}\right)\left[\frac{\left(R_{K, L}+R_{K-1, L}\right)}{2}-\frac{\left(R_{K, L-1}+R_{K-1, L-1}\right)}{2}\right] \\
-\Lambda_{K, L}\left(f_{K, L+1}-f_{K, L}\right)\left[\frac{\left(R_{K, L}+R_{K, L-1}\right)}{2}-\frac{\left(R_{K-1, L}+R_{K-1, L-1}\right)}{2}\right] \tag{2.20a}
\end{gather*}
$$

$(B f)_{K, L}$ is also a zone-centered quantity and combining Figs. 2.1a and 2.1b gives


There are three other equally valid choices for $\left(B^{i} f\right)$

$\left(B^{2} f\right)=k_{K, L}\left(R_{L}\right)_{K, L}-\ell_{K, L-1}\left(R_{K}\right)_{K, L}$


$$
\begin{equation*}
\left(B^{3} f\right)=k_{K-1, L}\left(R_{L}\right)_{K, L}-\ell_{K, L}\left(R_{K}\right)_{K, L} \tag{2.20c}
\end{equation*}
$$



$$
\begin{equation*}
\left(B^{4} f\right)=k_{K, L}\left(R_{L}\right)_{K, L}-\ell_{K, L-1}\left(R_{K}\right)_{K, L} \tag{2.20d}
\end{equation*}
$$

Since any of these is equally valid, an average of the four is taken

$$
\begin{equation*}
-\int R d R d Z D(\nabla f)^{2} \approx-\frac{1}{4} \sum_{K, L} \sum_{i=1}^{4}\left|B^{i} f\right|_{K, L}^{2}=\sum_{K, L} f_{K, L}(A f)_{K, L} V_{K, L} \tag{2.21}
\end{equation*}
$$

then

$$
\begin{align*}
& -\frac{1}{4} \sum_{K, L} \sum_{i=1}^{4}\left(B^{i} f\right)_{K, L}^{2}=\sum_{K, L} f_{K, L}(A f)_{K, L} V_{K, L}  \tag{2.22}\\
& -\frac{1}{4} \sum_{i=1}^{4}\left(B^{i} f\right)^{2}=f_{K, L}(A f)_{K, L} V_{K, L}
\end{align*}
$$

and finally,

$$
\begin{equation*}
-\frac{1}{4} \sum_{i=1}^{4}\left(B^{i}\right)^{T}\left(B^{i}\right)=V A \tag{2.23}
\end{equation*}
$$

Eq. (2.20a) and the counterparts for $\left(B^{2} F\right),\left(B^{3} f\right)$ and $\left(B^{4} f\right)$ are substituted into Eq. (2.23). On the left side, the coefficients of $f_{K, L} f_{K^{\prime}, L^{\prime}}$ are collected and equated with the coefficients of $f_{K, L} f_{K^{\prime}, L^{\prime}}$ on the right side to get

$$
\begin{gather*}
V_{K, L} A_{(K, L),(K, L)}=-\sigma_{(K, L)}-\sigma_{(K-1, L)}-\lambda_{(K, L)}-\lambda_{(K, L-1)} \\
+\frac{1}{2}\left(\rho_{(K, L)}^{1}+\rho_{(K, L)}^{2}-\rho_{(K+1, L)}^{3}-\rho_{(K, L)}^{4}\right)  \tag{2.24a}\\
V_{K, L} A_{(K, L),(K+1, L)}=\sigma_{(K, L)}-\frac{1}{4}\left(\rho_{(K, L)}^{1}+\rho_{(K+1, L)}^{2}-\rho_{(K+1, L)}^{3}-\rho_{(K, L)}^{4}\right)  \tag{2.24b}\\
V_{K, L} A_{(K, L),(K, L+1)}=\lambda_{(K, L)}-\frac{1}{4}\left(\rho_{(K, L)}^{1}+\rho_{(K, L+1)}^{2}-\rho_{(K, L)}^{3}-\rho_{(K, L+1)}^{4}\right)  \tag{2.24c}\\
V_{K, L} A_{(K, L),(K+1, L+1)}=-\frac{1}{4}\left(\rho_{(K+1, L)}^{3}+\rho_{(K, L+1)}^{4}\right)  \tag{2.24d}\\
V_{K, L} A_{(K, L),(K-1, L+1)}=\frac{1}{4}\left(\rho_{(K-1, L)}^{1}+\rho_{(K, L+1)}^{2}\right) \tag{2.24e}
\end{gather*}
$$

where

$$
\begin{align*}
\sigma_{(K, L)} & =\Sigma_{K, L}^{2} \frac{\left(R_{L}\right)_{K, L}^{2}+\left(R_{L}\right)_{K+1, L}^{2}}{2} \\
\lambda_{(K, L)} & =\Lambda_{K, L}^{2} \frac{\left(R_{K}\right)_{K, L}^{2}+\left(R_{K}\right)_{K, L+1}^{2}}{2} \\
\rho_{(K, L)}^{1} & =\Sigma_{(K, L)} \Lambda_{(K, L)} C_{(K, L)}  \tag{2.25}\\
\rho_{(K, L)}^{2} & =\Sigma_{(K-1, L)} \Lambda_{(K, L-1)} C_{(K, L)} \\
\rho_{(K, L)}^{3} & =\Sigma_{(K-1, L)} \Lambda_{(K, L)} C_{(K, L)} \\
\rho_{(K, L)}^{4} & =\Sigma_{(K, L)} \Lambda_{(K, L-1)} C_{(K, L)}
\end{align*}
$$

and

$$
\begin{equation*}
C_{K, L}=\left(R_{K}\right)_{K, L} \cdot\left(R_{L}\right)_{K, L} . \tag{2.26}
\end{equation*}
$$

Since $V A$ is symmetric

$$
\begin{equation*}
V_{K^{\prime}, L^{\prime}} A_{\left(K^{\prime}, L^{\prime}\right),(K, L)}=V_{K, L} A_{(K, L),\left(K^{\prime}, L^{\prime}\right)} \tag{2.27}
\end{equation*}
$$

and all elements of $A$ not defined by Eqs. 2.24a-e to 2.26 are zero.

### 2.3 Some derivations related to the code implementation

As shown before, the diffusion equation including the absorption and emission terms is:

$$
\begin{equation*}
\frac{\partial I}{c \partial t}-\nabla \cdot D \nabla I=4 \pi \sigma_{a}^{i} B_{v}-\sigma_{a}^{\prime} I . \tag{2.28}
\end{equation*}
$$

Using $\mathrm{I}=\mathrm{cE}$, and multiplying by the volume, the equation becomes

$$
\begin{equation*}
V \frac{\partial E}{\partial t}-V c \nabla \cdot D \nabla E=V 4 \pi \sigma_{a}^{\prime} B_{v}-V c \sigma_{a}^{\prime} E \tag{2.29}
\end{equation*}
$$

Using the implicit differencing scheme for the time variable,

$$
\begin{align*}
& V \frac{E^{n+1}-E^{n}}{\Delta t}-V c \nabla \cdot D \nabla E=V 4 \pi \sigma_{a}^{\prime} B_{v}-V c \sigma_{a}^{\prime} E^{n+1}  \tag{2.30}\\
& V\left(E^{n+1}-E^{n}\right)=\Delta t V c \nabla \cdot D \nabla E+\Delta t V 4 \pi \sigma_{a}^{\prime} B_{v}-\Delta t V c \sigma_{a}^{\prime} E^{n+1}
\end{align*}
$$

Replace $\nabla \cdot D \nabla E$ by AE as shown before,

$$
\begin{equation*}
V\left(E^{n+1}-E^{n}\right)=\Delta t V c A E E^{n+1}+\Delta t V 4 \pi \sigma_{a}^{\prime} B_{v}-\Delta t V c \sigma_{a}^{\prime} E^{n+1} \tag{2.31}
\end{equation*}
$$

Rearrange the above equation:

$$
\begin{equation*}
\left\{V-c \Delta t\left(V A-V \sigma_{a}^{\prime}\right)\right\} E^{n+1}=V E^{n}+\Delta t V 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.32}
\end{equation*}
$$

From the Kershaw scheme, we know the spatial differential matrix $\tilde{A}=V A$, and the volume for the last time step is $V^{n}$,

$$
\begin{equation*}
\left\{V^{n+1}-c \Delta t\left(\tilde{A}-V \sigma_{a}^{\prime}\right)\right\} E^{n+1}=V^{n} E^{n}+\Delta t V^{n+1} 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.33}
\end{equation*}
$$

Therefore, the diagonal elements and the nondiagonal elements are,

$$
\begin{aligned}
& \text { Diagonal }=V_{i, i}{ }^{n+1}-c \Delta t\left(\tilde{A}_{i, i}-V_{i, i} \sigma_{a}^{\prime}\right) \\
& \text { Nondiagonal }=-c \Delta t \tilde{A}_{i, j}
\end{aligned}
$$

Further define for convenience:

$$
\begin{equation*}
\overline{\bar{A}}=V^{n+1}-c \Delta t\left(\tilde{A}-V \sigma_{a}^{\prime}\right) \tag{2.34}
\end{equation*}
$$

The last equation becomes,

$$
\begin{equation*}
\overline{\bar{A}} E^{n+1}=V^{n} E^{n}+\Delta t V^{n+1} 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.35}
\end{equation*}
$$

Define $\Delta E^{n+1}=E^{n+1}-E^{n}$,

$$
\begin{equation*}
\overline{\bar{A}} \Delta E^{n+1}=V^{n} E^{n}-\overline{\bar{A}} E^{n}+\Delta t V^{n+1} 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.36}
\end{equation*}
$$

This is the final form that is solved by DRACO radiation diffusion transport module. As we can see, the actual result solved by the linear equation solver is the change of the radiation energy density, rather than the radiation energy density itself.

### 2.4 Variables in the code

In the situation that users want to understand how the scheme is actually implemented, we give brief descriptions about the variables in the code and their corresponding symbols in the equation. In the radiation module in the DRACO code, variable cmac_14 represents the diagonal elements of the matrix, and other 8 variables cmac_10-13, cmac_15-18 represent the nondiagonal
parts of the matrix. The code segment of implementation of the above equations in DRACO is listed as follows,

```
cmac_14(is:ie,js:je,1,igr) = cmac_14(is:ie,js:je,1,igr)
                - opacity_zone(is:ie,js:je,1,igr,iemiss)**ol(is:ie,js:je,1)
cmac_10(is:ie,js:je,1,igr) = cmac_10(is:ie,js:je,1,igr) *const1
cmac_11(is:ie,js:je,1,igr) = cmac_11(is:ie,js:je,1,igr) *const1
cmac_12(is:ie,js:je,1,igr) = cmac_12(is:ie,js:je,1,igr) *const1
cmac_13(is:ie,js:je,1,igr) = cmac_13(is:ie,js:je,1,igr) *const1
cmac_14(is:ie,js:je,1,igr) = cmac_14(is:ie,js:je,1,igr) *const1
cmac_15(is:ie,js:je,1,igr) = cmac_15(is:ie,js:je,1,igr) *const1
cmac_16(is:ie,js:je,1,igr) = cmac_16(is:ie,js:je,1,igr) *const1
cmac_17(is:ie,js:je,1,igr) = cmac_17(is:ie,js:je,1,igr) *const1
cmac_18(is:ie,js:je,1,igr) = cmac_18(is:ie,js:je,1,igr) *const1
cmac_rhs(is:ie,js:je,1,igr) = cmac_rhs(is:ie,js:je,1,igr) +
    radiation_energy_emission(is:ie,js:je,1,igr)*dt*vol(is:ie,js:je,1)
cmac_14(is:ie,js:je,1,igr) = vol(is:ie,js:je,1) - cmac_14(is:ie,js:je,1,igr)
cmac_rhs(is:ie,js:je,1,igr) = cmac_rhs(is:ie,js:je,1,igr) +
    vollast(is:ie,js:je,1)*radiation_energy_density(is:ie,js:je,1,igr)
cmac_rhs(is:ie,js:je,1,igr) = cmac_rhs(is:ie,js:je,1,igr)
    + cmac_10(is:ie,js:je,1,igr)*radiation_energy_density(is-1:ie-1,js-1:je-1,1,igr)
    + cmac_11(is:ie,js:je,1,igr)*radiation_energy_density(is :ie ,js-1:je-1,1,igr)
    + cmac_12(is:ie,js:je,1,igr)*}\mp@subsup{}{}{\star}radiation_energy_density(is+1:ie+1,js-1:je-1,1,igr)
    + cmac_13(is:ie,js:je,1,igr)*radiation_energy_density(is-1:ie-1,js :je ,1,igr)
    -cmac_14(is:ie,js:je,1,igr)*radiation_energy_density(is :ie ,js :je ,1,igr)
    + cmac_15(is:ie,js:je,1,igr)*radiation_energy_density(is+1:ie+1,js :je ,1,igr)
    + cmac_16(is:ie,js:je,1,igr)*radiation_energy_density(is-1:ie-1,js+1:je+1,1,igr)
    + cmac_17(is:ie,js:je,1,igr)*radiation_energy_density(is :ie ,js+1:je+1,1,igr)
    + cmac_18(is:ie,js:je,1,igr)*radiation_energy_density(is+1:ie+1,js+1:je+1,1,igr)
```

The variables in the code corresponding to the symbols in the equation are listed in the table,

| vol | $V_{i, i}{ }^{n+1}$ |
| :--- | :--- |
| const1 | $c \Delta t$ |
| vollast | $V_{i, i}{ }^{n}$ |
| opacity_zone | $\sigma_{a}^{\prime}$ |
| radiation_energy_emission | $4 \pi \sigma_{a}^{\prime} B_{v}$ |
| cmac | $\overline{\bar{A}}$ |

We also want to point out the difference between the current implementation of the above eaquation and Verdon's implementation. (We don't know exactly Verdon's scheme. We can only tell the slight implementation difference from the code segment.)

The Verdon's code segment is:

```
cmac_14(i,j,1,igr) = dt*cmac_14(i,j,,1,igr)-(const3*alpha)-(const4*alpha)
cmac_14(i,j, ,1,igr) = cmac_14(i,j,1,igr)*vol(i,j,1)
cmac_17(i,j,1,igr) = cmac_17(i,j,1,igr)*const5
cmac_18(i,j,1,igr) = cmac_18(i,j,1,igr)*const5
cmac_15(i,j,1,igr) = cmac_15(i,j,1,igr)*const5
cmac_12(i,j,1,igr) = cmac_12(i,j,1,igr)*const5
cmac_11(i,j,1,igr) = cmac_11(i,j,1,igr)*const5
cmac_10(i,j,1,igr) = cmac_10(i,j,1,igr)*const5
cmac_13(i,j,1,igr) = cmac_13(i,j,1,igr)*const5
cmac_16(i,j,1,igr) = cmac_16(i,j,1,igr)*const5
cmac_rhs(i,j,1,igr) = radiation_energy_emission(i,j,1,igr)*const5
cmac_rhs(i,j,1,igr) = cmac_rhs(i,j,1,igr) +
    ((cmac_14(i,j,1,igr)*radiation_energy_density(i ,j ,1,igr) +
    cmac_17(i,j,1,igr)*radiation_energy_density(i ,j+1,1,igr) +
    cmac_18(i,j,1,igr)*radiation_energy_density(i+1,j+1,1,igr) +
    cmac_15(i,j,1,igr)*radiation_energy_density(i+1,j ,1,igr) +
    cmac_12(i,j,1,igr)*radiation_energy_density(i+1,j-1,1,igr) +
    cmac_11(i,j,1,igr)*radiation_energy_density(i ,j-1,1,igr) +
    cmac_10(i,j,1,igr)*radiation_energy_density(i-1,j-1,1,igr) +
    cmac_13(i,j,1,igr)*radiation_energy_density(i-1,j ,1,igr) +
    cmac_16(i,j,, 1,igr)*radiation_energy_density(i-1,j+1,1,igr)) * rtalphai)
    cmac_14(i,j,,1,igr) = vollast(i,j,1)-cmac_14(i,j,1,igr)
```

From this code, we guess Verdon followed the derivation like this:
From Equation (2.32):

$$
\begin{equation*}
\left\{V^{n+1}-c \Delta t\left(\tilde{A}-V \sigma_{a}^{\prime}\right)\right\} E^{n+1}=V^{n} E^{n}+\Delta t V^{n+1} 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.37}
\end{equation*}
$$

Write $c \Delta t\left(\tilde{A}-V \sigma_{a}^{\prime}\right)$ as $\vec{A}$ (we use $\vec{A}$ to distinguish from the above matrix symbol),

$$
\begin{equation*}
V^{n} E^{n+1}-\stackrel{\rightharpoonup}{A} E^{n+1}=V^{n} E^{n}+\Delta t V^{n+1} 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.38}
\end{equation*}
$$

Note $V^{n+1}$ in Equation (2.37) changes to $V^{n}$ in Equation (2.38), and we also define $\Delta E^{n+1}=E^{n+1}-E^{n}$,

$$
\begin{equation*}
\left(V^{n}-\vec{A}\right) \Delta E^{n+1}=\vec{A} E^{n}+\Delta t V^{n+1} 4 \pi \sigma_{a}^{\prime} B_{v} \tag{2.39}
\end{equation*}
$$

Therefore, the difference comes from the treatment of the volume and the definition of the intermediate matrix $\vec{A}$.

### 2.5 Boundary Conditions

The logical mesh consists of two types of zones: physical and vacuum or ghost zones. The vacuum zones completely surround the physical zones. With the use of vacuum zones, it is possible to have irregular-shaped physical zone boundaries while still maintaining the rectangular logical mesh (i.e., for a given $K=1, \ldots, K M A X$ and $L=1, \ldots, L M A X, K M A X$ and $L M A X$ are constant throughout the mesh).

The two types of boundary conditions considered are the Dirichlet condition or "escape" boundary

$$
f=f_{B}
$$

where $f_{B}$ is a specified flux in the vacuum, and the reflective or "no escape" boundary

$$
d S \cdot \nabla f=0
$$

To implement these boundary conditions, it is necessary to revise the definitions of $\Sigma_{K, L}$ and $\Lambda_{K, L}$ on the boundaries. The revisions depend upon the type of zone it is and whether the neighboring zones are vacuum or physical zones. The following rules are used to determine the components $\Sigma_{K, L}$ and $\Lambda_{K, L}$. The rules are defined for $\Sigma_{K, L}$, which connects the zone $(K, L)$ with the zone $(K+1, L)$. An analogous sent of rules applies for $\Lambda_{K, L}$, which connects the zone $(K, L)$ with the zone ( $K, L+1$ ).

1) If ( $K, L$ ) and ( $K+1, L$ ) are both physical zones $\Sigma_{K, L}\left(R_{L}\right)_{K, L}$ and $\Sigma_{K, L}\left(R_{L}\right)_{K+1, L}$ are defined as previously (using Eqs. 2.16). This is corresponding to the physical region for the problem.
2) If ( $K, L$ ) and ( $K+1, L$ ) are both vacuum zones, $\Sigma_{K, L}=0$.
3) If $(K, L)$ is a physical zone and $(K+1, L)$ is a vacuum zone and the face separating them has a "no escape" boundary condition, $n \cdot \nabla f=0$, where n is the unit normal to the vacuum surface then $\Sigma_{K, L}=0$ and $\left(R_{K}\right)_{K, L}$ is replaced by

$$
n\left(n \cdot\left(R_{K}\right)_{K, L}\right)
$$

4) If ( $K, L$ ) is a physical zone and $(K+1, L)$ is a vacuum zone and the face separating them has an "escape" boundary condition, then $\Sigma_{K, L}\left(R_{L}\right)_{K, L}$ remains defined as previously (Eqs. 2.6), but now in the expression

$$
\Sigma_{K, L}\left(R_{L}\right)_{K+1, L},
$$

we use

$$
\left(R_{L}\right)_{K+1, L}=R_{K, L}-R_{K, L-1},
$$

since $R_{K+l, L}$ and $R_{K+l, L-l}$ do not exist, and $\left(R_{K}\right)_{K, L}$ is replaced by

$$
n\left(n \cdot\left(R_{K}\right)_{K, L}\right)
$$

n is the unit normal to the vacuum surface. $\Sigma_{K, L}$ is now $(D R / j)^{1 / 2}$ suitably averaged between zone $(K, L)$ and the vacuum, but since $D$ is not defined in the vacuum zone, $\Sigma_{K, L}$ is defined by

$$
\Sigma_{K, L}^{2}=\left(\frac{D_{K, L} R_{K, L}}{j_{K, L}}\right) \cdot \frac{1}{2}
$$

The above cases (3) and (4) correspond to the right-side boundary for the problem.
5) If ( $K, L$ ) is a vacuum zone and $(K+1, L)$ is a physical zone and the face separating them has a "no escape" boundary condition, $n \cdot \nabla f=0$, then $\Sigma_{K, L}=0$ and $\left(R_{K}\right)_{K+1, L}$ is replaced by

$$
n\left(n \cdot\left(R_{K}\right)_{K+1, L}\right)
$$

6) If ( $K, L$ ) is a vacuum zone and $(K+1, L)$ is a physical zone and the face separating them has an "escape" boundary condition, $f=f_{B}$, then in the expression

$$
\left.\Sigma_{K, L}\left(R_{L}\right)_{K, L}\right),
$$

we use

$$
\left(R_{L}\right)_{K, L}=R_{K, L}-R_{K, L-1}
$$

$\left(R_{K}\right)_{K, L}$ is replaced by

$$
n\left(n \cdot\left(R_{K}\right)_{K+1, L}\right)
$$

The above cases (5) and (6) correspond to the left-side boundary for the problem.
7) If $(K, L)$ is a physical zone and $(K, L+1)$ is a vacuum zone and the face separating them has a "no escape" boundary condition, $n \cdot \nabla f=0$, where n is the unit normal to the vacuum surface then $\Lambda_{K, L}=0$ and $\left(R_{L}\right)_{K, L}$ is replaced by

$$
n\left(n \cdot\left(R_{L}\right)_{K, L}\right)
$$

8) If ( $K, L$ ) is a physical zone and $(K, L+1)$ is a vacuum zone and the face separating them has an "escape" boundary condition, then $\Lambda_{K, L}\left(R_{K}\right)_{K, L}$ remains defined as previously (Eqs. 2.6), but now in the expression

$$
\Lambda_{K, L}\left(R_{K}\right)_{K, L+1},
$$

we use

$$
\left(R_{K}\right)_{K, L+l}=R_{K, L}-R_{K-l, L},
$$

since $R_{K, L+1}$ and $R_{K-1, L+1}$ do not exist, and $\left(R_{L}\right)_{K, L}$ is replaced by

$$
n\left(n \cdot\left(R_{L}\right)_{K, L}\right)
$$

n is the unit normal to the vacuum surface. $\Lambda_{\mathrm{K}, \mathrm{L}}$ is now $(D R / j)^{1 / 2}$ suitably averaged between zone ( $K, L$ ) and the vacuum, but since $D$ is not defined in the vacuum zone, $\Lambda_{K, L}$ is defined by

$$
\Lambda_{K, L}^{2}=\left(\frac{D_{K, L} R_{K, L}}{j_{K, L}}\right) \cdot \frac{1}{2}
$$

The above cases (7) and (8) correspond to the top-side boundary for the problem.
9) If ( $K, L$ ) is a vacuum zone and $(K, L+1)$ is a physical zone and the face separating them has a "no escape" boundary condition, $n \cdot \nabla f=0$, then $\Lambda_{K, L}=0$ and $\left(R_{L}\right)_{K, L+1}$ is replaced by

$$
n\left(n \cdot\left(R_{L}\right)_{K, L+1}\right)
$$

10) If ( $K, L$ ) is a vacuum zone and $(K, L+1)$ is a physical zone and the face separating them has an "escape" boundary condition, $f=f_{B}$, then in the expression

$$
\left.\Lambda_{K, L}\left(R_{L}\right)_{K, L}\right)
$$

we use

$$
\left(R_{L}\right)_{K, L}=R_{K, L}-R_{K, L-1}
$$

and $\left(R_{L}\right)_{K, L}$ is replaced by

$$
n\left(n \cdot\left(R_{L}\right)_{K, L}\right)
$$

The above cases (9) and (10) correspond to the bottom-side boundary for the problem.

In summary, the treatment of the nonescape boundary condition is much easier than the free surface boundary condition since we know the values at the boundary because of the reflection and adiabatic property of the boundary. However, for the free escape boundary condition, we don't know the exact solution at the boundary. The best we can do is assume at some extrapolated distance the value of f goes to zero, as described in Milne's problem [5]. In DRACO, we use this approximation. We use the right-side boundary as an example (Fig. 2.2)


Fig. 2.2. Milne right-side boundary condition.
$\mathrm{E}=0.71 \lambda$, where $\lambda$ is the transport mean free path. In addition, the calculation of $\Sigma_{K, L}$ (Eq. 2.16) which is averaged between the last zone and the vacuum zone is changed to include the area defined by the extrapolated length as follows

$$
\Sigma_{K, L}^{2}=\frac{R_{K, L} \cdot D_{K, L}}{A_{K, L}}
$$

and the Jacobian area $A_{K, L}=j_{K, L}+A_{E x}$, where $A_{E x}$ is the Jacobian area determined by the extrapolated length.

### 2.6 Flux limiter

When dealing with the moment of the transport equation, there always appears one unknown quantity. To close the set of equations, an approximation is needed to break the chain [5]. The P1 equation is obtained by setting the radiation pressure equal to one-third of the energy density. It has first-order accuracy in the optically thick limit, while in the optically thin limit it has an incorrect propagation velocity. The diffusion equation makes a further approximation to the P1 equation by omitting the gradient of the flux such that the cutoff equation is in the form of the

Fick's law. Morel [6] has shown in the diffusive limit the diffusion equation has the first-order accuracy comparable with the P1 equation. However, in the optically thin limit, this equation has an infinite propagation velocity. To correct this problem, the flux limiter is always introduced.

In DRACO, four options are available [7] beside one option (flux_limiter_type=0) that has no flux limiter included.
a) Option 1

Option 1 uses the summation form of the transport cross section and the gradient. It is called the "sum" flux limiter.

$$
D=\frac{1}{3 \sigma_{t r}+\frac{\delta}{E}|\nabla E|} .
$$

The factor $\delta$ is an adjustable parameter for which we set a default value of one. In the optically thick limit, the gradient of the energy density is small compared to the transport mean free path and therefore the equation tends to be diffusive. In the thin limit, the gradient dominates the denominator of the diffusion coefficient and the flow is limited to the speed of light, as desired. The introduction of the gradient leads the equation to be nonlinear in the energy density. Morel has shown this form is only zero order accurate in the thick limit but gets the correct approximation in the optically thin limit.
b) Option 2

Larsen [7] has suggested a modified form which retains the first-order accuracy in the optically thick limit:

$$
D=\left[\frac{1}{\left(3 \sigma_{t r}\right)^{n}+\left(\frac{\delta}{E}|\nabla E|\right)^{n}}\right]^{1 / n}
$$

In DRACO, we use $\mathrm{n}=2$.

## c) Option 3

The maximum of the transport cross section and the gradient is used. It is called the "max" flux limiter. The form also retains the first-order accuracy in the thick limit when the cross section is larger than the gradient term. However, this flux limiter has discontinuous derivatives.

$$
D=\frac{1}{\max \left(3 \sigma_{t r}, \frac{\delta}{E}|\nabla E|\right)}
$$

d) Option 4

Another flux limiter is proposed by Levermore and Pomraning (LP) [8] by solving exactly a particular transport problem. Olson has shown this flux limiter exhibits some nonphysical behavior. A simplified form of this kind of limiter is suggested by Zimmerman:

$$
\begin{gathered}
D=\frac{\xi(\beta)}{\sigma_{t r}} \\
\xi(\beta)=\left[\operatorname{coth}(\beta)-\frac{1}{\beta}\right] \frac{1}{\beta} \\
\beta=\frac{|\nabla E|}{\sigma_{t r} E}
\end{gathered}
$$

These options can be selected by setting the "flux_limiter_type" in the input parameters.

## 3. Test problems

We have done a series of tests to look at the correctness of the implementation by running the code in different geometries and different meshes. Two kinds of meshes are used in the test problem as shown in Fig. 3.1 and Fig. 3.2. Both meshes are also applied to the planar and cylindrical geometries. All of the simple test problems were run with graphics. Figures 3.3-3.17 illustrate the time evolution of the radiation energy under various boundary conditions and meshes. The graphical results are easily understandable with the brief graphic description.

To test the accuracy of the scheme, we use three simple problems that have an exact analytical solution. Figure 3.18 shows the result for the problem that has the left boundary condition with the energy density equal to zero and the right boundary condition with the energy
density equal to 8 . There is an exact solution for this problem in one dimension, that is, $\mathrm{f}=8 \mathrm{Z}$. We simulate the one dimension by setting the reflective boundary condition for both top and bottom boundaries. We can see the numerical result agrees with the exact solution very well. The second problem we solved is the wave propagation with a delta function source in the middle of the onedimensional plane. To be more specific, the diffusion equation under this case is [9]

$$
\frac{\partial T}{\partial t}=\frac{\partial}{\partial x} D \frac{\partial T}{\partial x}
$$

with the initial condition of $\mathrm{T}(\mathrm{x} / 2,0)=\mathrm{Q} \delta(\mathrm{x} / 2)$. The solution to the problem for the constant diffusion coefficient is

$$
T=\frac{Q}{\sqrt{4 \pi D t}} \exp \left(-\frac{\left(x-x_{0} / 2\right)^{2}}{4 D t}\right) .
$$

The simulation starts with a distribution of the above form at some initial time. Figure 3.19 shows the comparison of the exact solution and the numerical result. While Fig. 3.19 shows the result for the linear conduction, Fig. 3.20 shows the result for the nonlinear wave conduction. The diffusion equation for this case is

$$
\frac{\partial T}{\partial t}=a \frac{\partial}{\partial x} T^{5} \frac{\partial T}{\partial x}
$$

The diffusion equation is nonlinear because the diffusion coefficient is a function of T . The coefficient has also very large gradient since it is a function of T to the fifth power. Again, the delta source Q is added at the middle of the plane at the initial time. The exact solution for this case [9] is

$$
\begin{aligned}
& T=1.12 \frac{Q}{2 x_{f}}\left(1-\frac{x^{2}}{x_{f}^{2}}\right)^{1 / 5} \\
& x_{f}=0.77\left(a Q^{n} t\right)^{\frac{1}{(n+2)}} .
\end{aligned}
$$

We compare the numerical result with the exact solution in Fig. 3.20. Again, they agree very well except that there is a little discrepancy at the wave front. This results from the average treatment of the diffusion coefficient at the cell face center, as in Eq. (2.16).


## Notation used in the figure captions:

| Omesh: | orthogonal mesh as Fig. 3.1 |
| :--- | :--- |
| Zmesh: | distorted mesh as Fig. 3.2 |
| Planar: | planar geometry |
| Cyd: | cylindrical geometry |
| R: | right |
| L: | left |
| T: | top |
| B: | bottom |
| Escape bc: | escape boundary condition $(\mathrm{f}=0)$, |
|  | otherwise, adiabatic boundary condition. |



Fig. 3.3. Omesh, planar, L\&R escape bc.


Fig. 3.5. Omesh, planar, T\&B escape bc.


Fig. 3.7. Omesh, planar, T\&B\&L\&R escape bc.


Fig. 3.4. Zmesh, planar, L\&R escape bc.


Fig. 3.6. Zmesh, planar, T\&B escape bc.


Fig. 3.8. Zmesh, planar, T\&B\&L\&R escape bc.


Fig. 3.9. Omesh, cyd, T\&R escape bc.


Fig. 3.11. Omesh, cyd, T\&L escape bc.


Fig. 3.13. Omesh, cyd, T\&L\&R escape bc.


Fig. 3.10. Zmesh, cyd, T\&R escape bc.


Fig. 3.12. Zmesh, cyd, T\&L escape bc.


Fig. 3.14. Zmesh, cyd, T\&L\&R escape bc.


Fig. 3.15. Zmesh, cyd, f=8(T).


Fig. 3.17. Omesh, planar, $f=0(L) f=8(R)$.


Fig. 3.19. Omesh, planar, Gaussian distribution test.


Fig. 3.16. Zmesh, planar, $f=0(B) f=8(T)$.


Fig. 3.18. Comparison with exact solution for Fig. 3.17.


Fig. 3.20. Omesh, planar, nonlinear diffusion coef test.

## 4. Subroutines and namelist options

This section is devoted to describe the subroutines related to the radiation transport module in DRACO and the namelist options in the input file. We list the subroutines and their purpose in Table 4.1.

Figures 4.1-4 show the flow chart of radiation transport.
If only the radiation emission is considered, no radiation transport will be carried out. The energy contribution for electron source from radiation is simply from the radiation emission. Two diffusion schemes are available for building the 9-point diffusion matrix, that is, the Kershaw scheme and the Verdon scheme. After the matrix is set up, it is solved by the same solver routine "math_solver_rad_trans_2d".

Table 4.1 Subroutines in the radiation transport module

| rad_trans_control | controls the radiation transport routines |
| :---: | :---: |
| rad_trans_emission_lte | determine the frequency dependent emission |
| rad_trans_opac_offtable_lte | determine the frequency dependent local thermal equilibrium opacities |
| rad_trans_opac_table_le | get frequency dependent local thermal equilibrium opacities via table lookup |
| rad_trans_elec_srce_trms_2d | determines the electron source terms determined by the change of in the radiation field |
| elec_srce_trms_2d_no_transport | determines the electron source terms determined by the total emission |
| rad_gradient_2d | calculate the radiation energy gradient |
| math_solver_rad_trans_2d | solve for the change in radiation energy density with multiple solver options |
| rad_trans_coefs_2d | determines the matrix for the radiation diffusion under the old Verdon's scheme |
| rad_trans_coefs_2d_kershaw | determines the matrix for the radiation diffusion under the Kershaw's scheme |



Fig. 4.1. Main control of the radiation transport.


Fig. 4.2 b

Fig. 4.2a. Flow chart of Kershaw scheme.


Fig. 4.2b. Flow chart of Kershaw scheme.


Fig. 4.3. Flow chart of math solver for the diffusion differencing linear equation.

```
elec_srce_trms_2d_no_transport
```

1) calc ulate electron energy deposition term by Setting rad_elec_src_term equal to the summation Of radiation_energy_emission for all groups
2) Time steps for radition
```
rad_trans_elec_srce_trms_2d
```

1) Calculate the radiation pressure
2) Calculate the rad_elec_src_term term by taking account Of the absorption term: absorption term-emission term
3) Time steps for radition

Fig. 4.4. Coupling of radiation and plasma. The radiation contributes the pressure to the hydrodynamics and energy (only emission energy or with the absorption energy) to the electron energy source term.

The namelist options are listed in Table 4.2

| Variables | Def. value | Dim. | Meaning |
| :---: | :---: | :---: | :---: |
| opacity_table_format | lle | s | Opacity table format. Three options are available: "Wisconsin", "TOPS", "lle". |
| rad_trans_bc_ism | zero | s | Left-side boundary condition. Two options are available: "zero"-no escape, reflective; "none"free boundary, escape. |
| rad_trans_bc_iep | none | S | Right-side boundary condition. Options as in rad_trans_bc_ism. |
| rad_trans_bc_jsm | zero | s | Bottom-side boundary condition. Options as in rad_trans_bc_ism. |
| rad_trans_bc_jep | zero | S | Top-side boundary condition. Options as in rad_trans_bc_ism. |
| rad_trans_bc_ksm | zero | S | For three dimension problem. Not implemented yet. |
| rad_trans_bc_kep | zero | S | For three dimension problem. <br> Not implemented yet. |
| number_of_radiation _freq_groups | 0 | S | Number of radiation frequency groups. |
| solver_option_rt | iccg | S | Choice of radiation diffusion matrix solver. Four options are available: <br> "yale" - Yale direct solver; <br> "iccg" - iccg iterative solver; <br> "petsc" - parallel PETSC library [10] solver across all processors; <br> "petscd" - parallel PETSC library solver groupwised. |
| petsc_ksp_type_rt | preonly | S | Krylov Subspace method choice. Twelve options are available: <br> 'richardson' - Richardson method; 'chebychev' - Chebychev method; 'cg' - Conjugate Gradient method; 'bicg' - BiConjugate Gradient method; 'gmres' - Gen. Min. Residual method; 'bcgs' - BiCGSTAB method; 'cgs' - Con. Gradient sq. method; 'tfqmr' - Trans-free QMR 1 method; 'tcqmr' - Trans-free QMR 2 method; 'cr' - Conjugate Residual method; |


|  |  |  | 'lsqr' - Least Squares method; 'preonly' - Shell for no KSP method. |
| :---: | :---: | :---: | :---: |
| petsc_pc_type_rt | lu | S | Choice of preconditioner types. Thirteen options are available: <br> 'jacobi' - Jacobi precondictioner; <br> 'bjacobi' - Block Jacobi preconditioner; 'sor' - SOR method; 'eisenstat' - SOR w/Eisenstat preconditioner; 'icc' - Incomplete Cholesky preconditioner; 'ilu' - Incomplete LU preconditioner; 'asm' - Additive Schwarz preconditioner; 'sles' - Linear solver preconditioner; 'composite' - Combin. preconditioners; 'lu' - LU preconditioner; 'cholesky' - Cholesky preconditioner; 'none' - no preconditioner; 'shell' - Shell for user defined preconditioner. |
| radiation_schema | kershaw | S | Verdon's diffusion matrix scheme. <br> Kershaw's diffusion matrix scheme. |
| verbose_rt | false | S | Prints solver iterations each time step. |
| mprtfrt | 99999 | S | Radiation transport information print frequency. |
| number_iters_limit_ solver_rt | 1000 | S | Maximum number of iterations for iterative solver. |
| number_iters_max_ solver_rt | 1000 | S | Maximum number iterations per cycle for radiation transport solver. |
| tot_its_group_rt | 1000 | 1 | total iterations per group for radiation. |
| solver_atol_rt | 10 | S | absolute size of residual norm for rt solver |
| solver_dtol_rt | 1.0 E 100 | S | Relative increase in the residual. |
| solver_rtol_rt | $1.0 \mathrm{E}-06$ | S | Relative converge test. |
| $\begin{aligned} & \text { rad_trans_energy_ } \\ & \text { limit } \end{aligned}$ | $1.0 \mathrm{E}-10$ | S | radiation transport energy limit, if less than this value, no radiation transport. |
| flux_limiter_type | 1 | S | radiation transport flux_limiter_type. Five options are available: <br> 0 - no limiter; <br> 1 - sum limiter; <br> 2 - quadrilateral limiter; <br> 3 - max limiter; 4 - LP limiter. |
| emission_only | false | S | Only sink term for radiation. No radiation transport. |

s denotes scalar value.

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