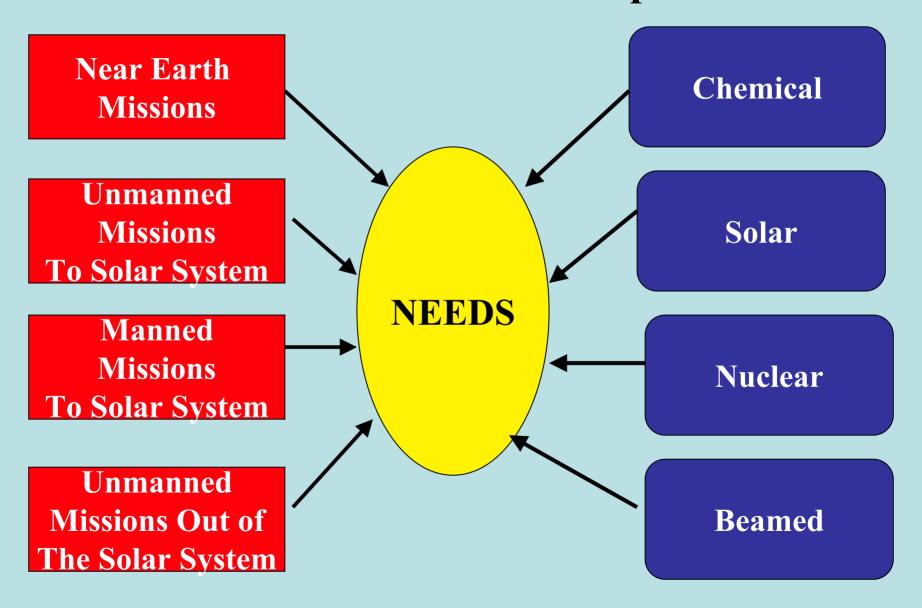
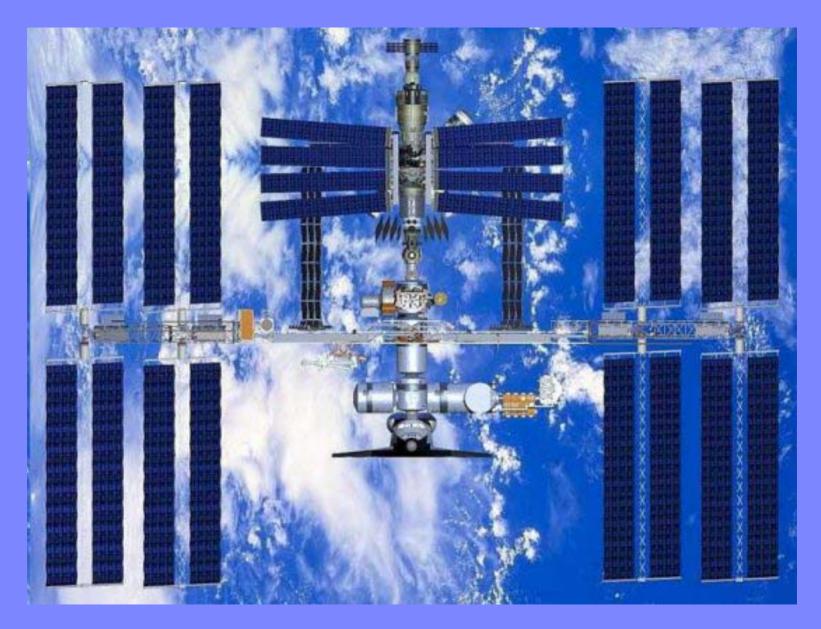


There are Many Requirements and Solutions to the Power Needs in Space

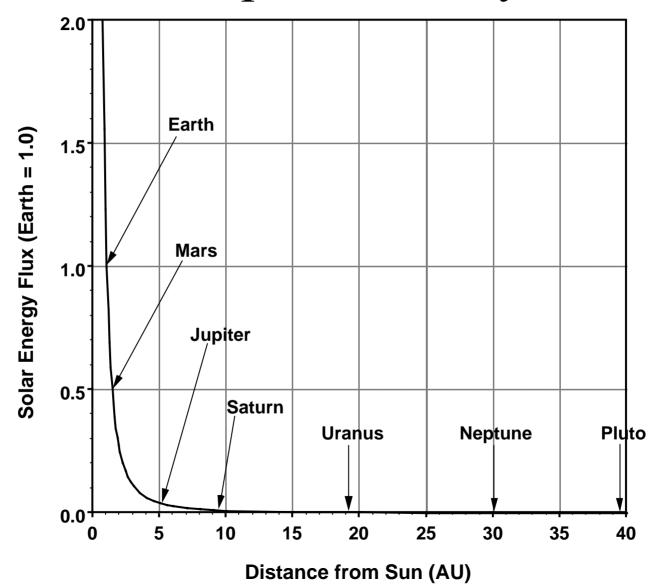




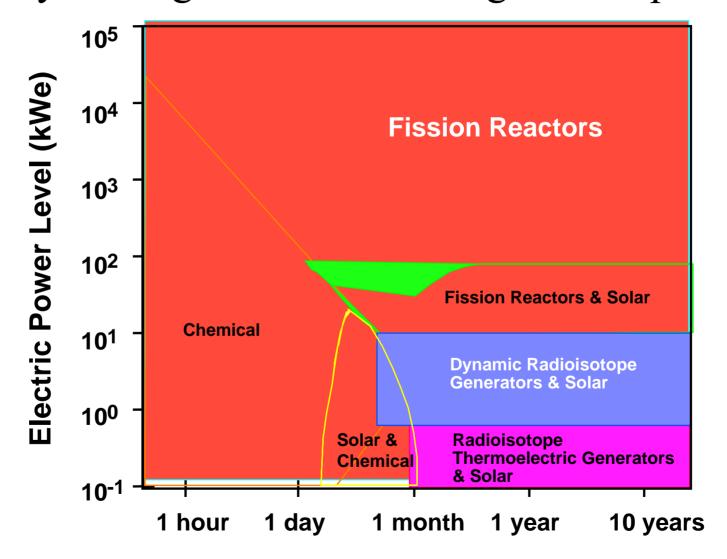
International Space Station Power Requirements -Full Station Envisioned-

- Total continuous needs->105 kW_e (MIR \approx 30kW_e)
- Two independent PV supplies (US=79 kW_e, Russian=29 kW_e)
- 120 VDC for US, 28 VDC for Russian system
- US array-33.1 by 73.2 meters (54% of football field)
- Mass $\approx 3.1 \text{ kg/m}^2$, 7.56 tonnes
- Power density $\approx 100 \text{ kg/kW}_e$
- US system-24 NiH batteries (eclipse, 168 kg, 6.5 y)
- Plus coolant to keep batteries @ 0-10 °C

Solar Power is Impractical Beyond Mars



The Use of Nuclear Power in Space is Absolutely Necessary for High Power and Long Time Operations

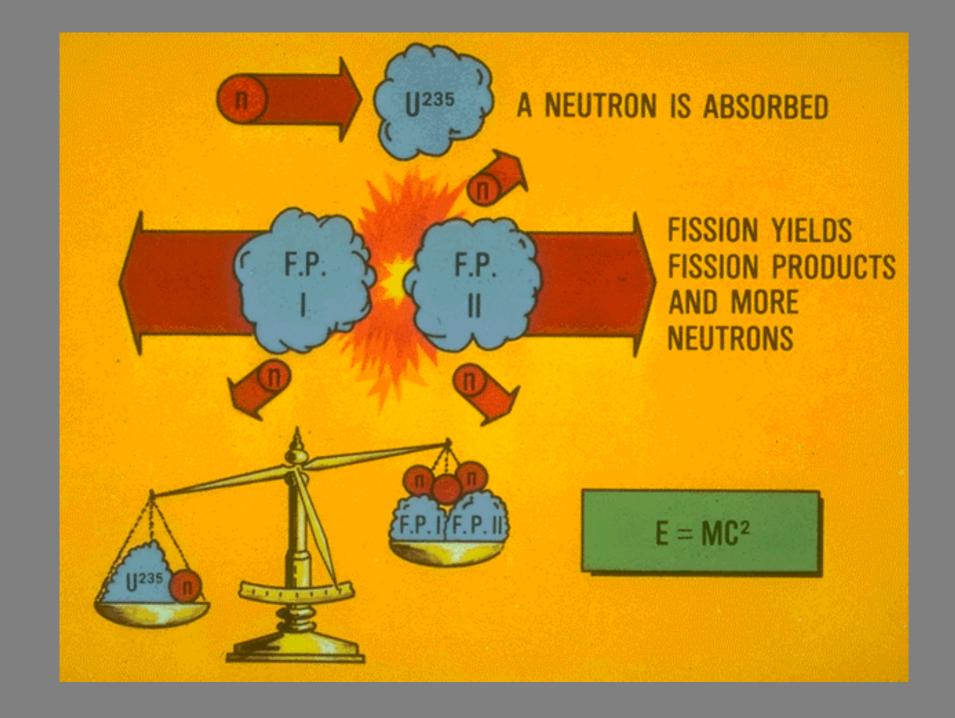


What is the Advantage of Using Nuclear Energy in Space?

1 kg of nuclear fuel contains

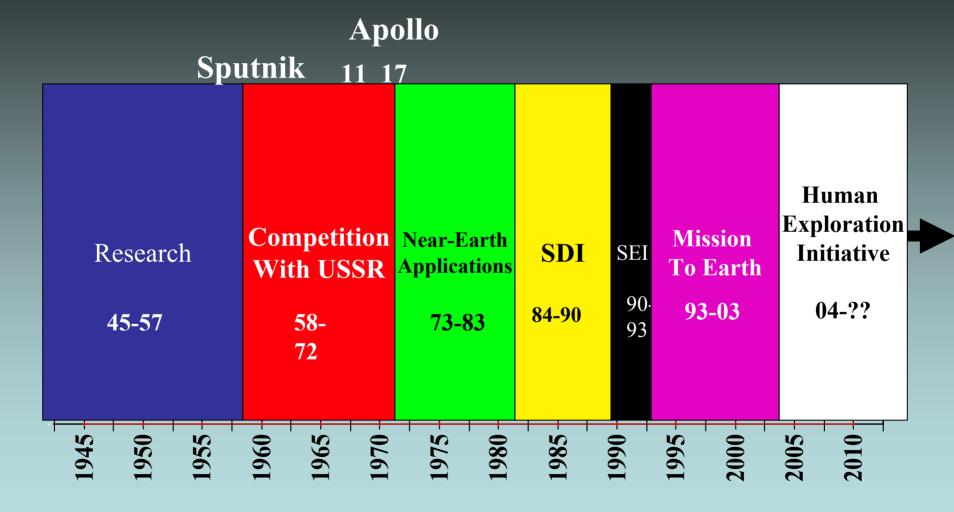
10,000,000

times the energy in 1 kg of chemicals

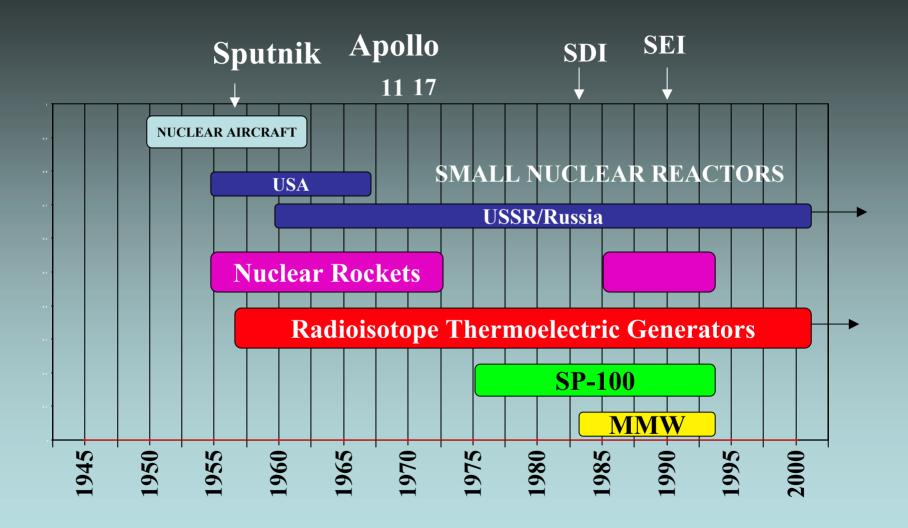


Nuclear Energy Can Be Converted to Electricity in a Variety of Ways **Nuclear Heat Source** Radioisotopes **Nuclear Reactors** Static **Dynamic** Rankine **Thermoelectrics Brayton Thermionics Stirling**

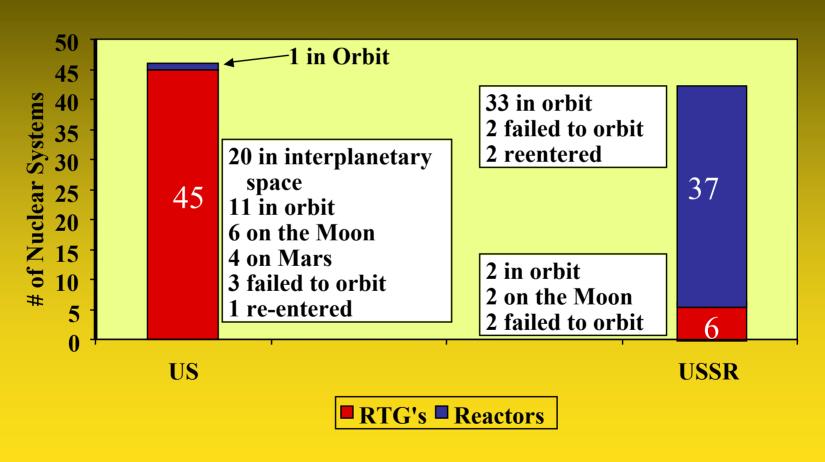
There Have Been Many Driving Forces Behind the Development of Nuclear Power in Space



CHRONOLOGY OF SPACE NUCLEAR POWER DEVELOPMENT



The U. S. and USSR Took Different Approaches to Nuclear Power Units in Space



RTG's Produce Power by Radioactive Decay

- Half life of radioactive species:
 - Law of Radioactive Decay $dN/dt = -\lambda N$
 - Integrating, $N(t) = N_o \exp(-\lambda t)$
 - Where the decay constant $\lambda = \ln 2/t_{1/2} = 0.693/t_{1/2}$
- Units of radioactive decay rate:
 - -1 Curie = 3.7 x 10^{10} disintegrations/s (dps)
 - -1 Becqeral = 1 dps

The Energy Released Depends on the Mass Difference Between Isotopes

• Maximum energy released in the decay of parent isotope

$$\begin{array}{c} {}^{A}X_{z} \text{ to daughter's } \sum {}^{A}{}_{i}Y_{Zi} \\ \\ E = \Delta m \cdot c^{2} \\ E = (mass \, {}^{A}X_{z} - mass \, \sum {}^{A}{}_{i}Y_{Zi}) \cdot c^{2} \end{array}$$

• Example:

238
Pu₉₄ --> 234 U₉₂ + 4 He₂ 1 amu=931.5 MeV
Mass 238 Pu₉₄ =238.049555 amu

Mass $^{234}U_{92} = 234.040947$ amu

Mass ${}^{4}\text{He}_{2} = 4.002603$ amu

 $\Delta m = 0.006005 \text{ amu}$ or E = 5.59 MeV

Power Generated by Radioisotopic Heat Sources

Initial # of Radioisotopes

N_o

Decay Rate, s-1

$$\lambda = \frac{\ln 2}{t_{1/2}}$$

Energy Released Per Decay MeV E

Number of Radioisotopes Left After t Years

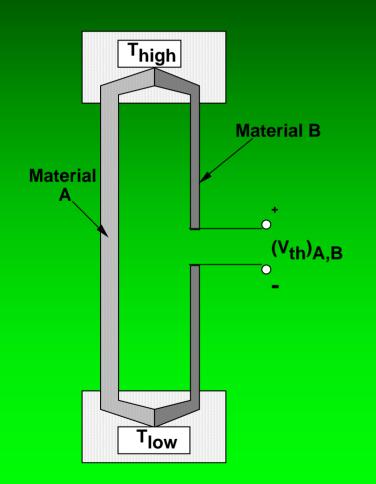
$$\exp -\left(\frac{t^*\ln 2}{t_{1/2}}\right)$$

Power (t) =
$$\frac{N_0 * E * ln 2}{t_{1/2}}$$
 exp -($\frac{t* ln 2}{t_{1/2}}$)

²³⁸Pu-The Radioisotope of Choice for Long Term Space Missions

- Half life-87.4 years
- Energy released per decay-5.6 MeV
- Specific activity-17 Ci/g
- Specific power density-30 Ci/W
- Power density-0.56 W/g
- Energy content for 10y mission-47 kWh/g
- Useful form-PuO₂ (MP = 2,250 °C)
- Production rate in fission reactor:
 - $-15 \text{ kg/1,000 MW}_{e}\text{y}$
- Cost of ²³⁸Pu-\$300/g

Thermoelectricity-A Reliable Way to Convert Heat Energy Directly into Electricity

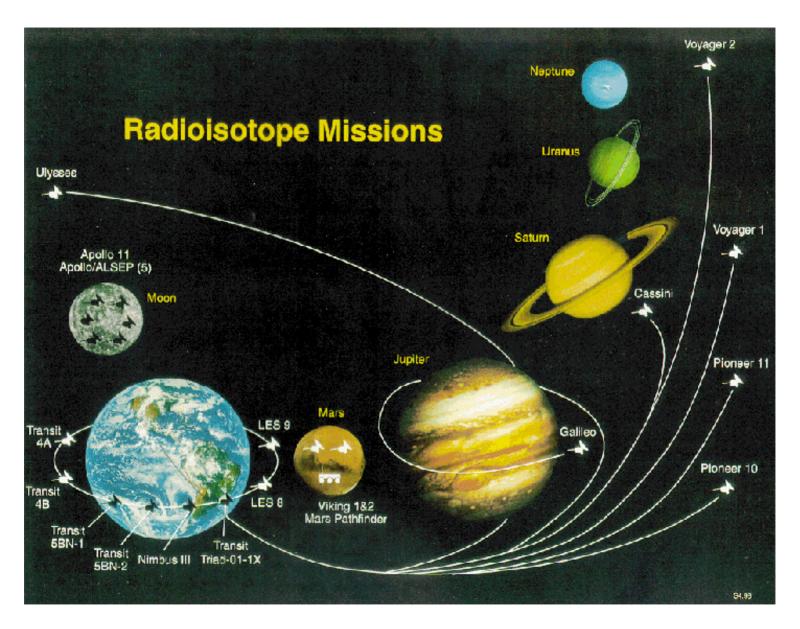


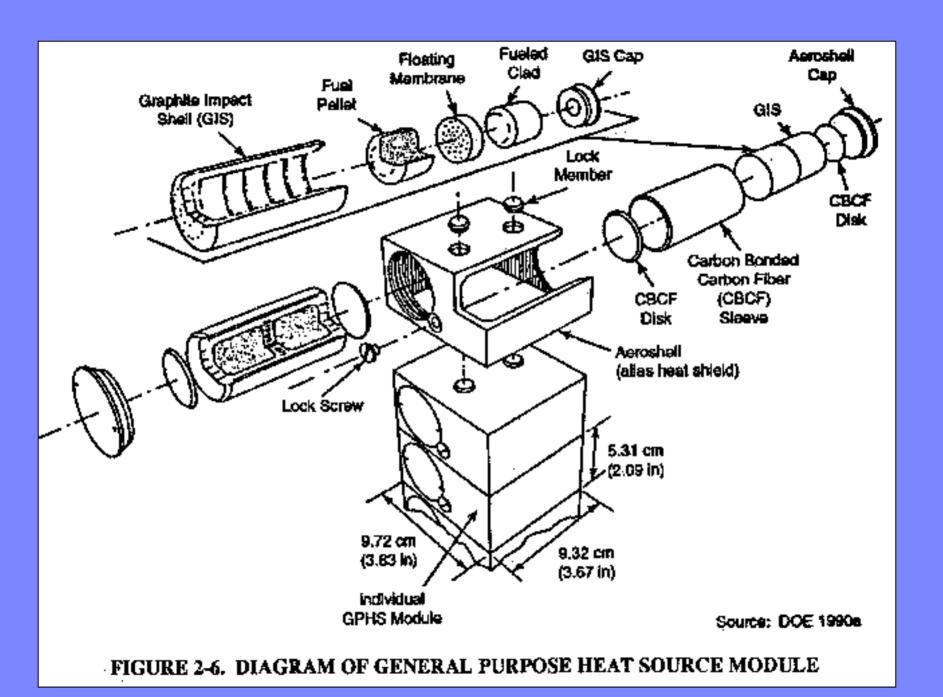
efficiency = $\eta_{carnot} \cdot \eta_{mat}$

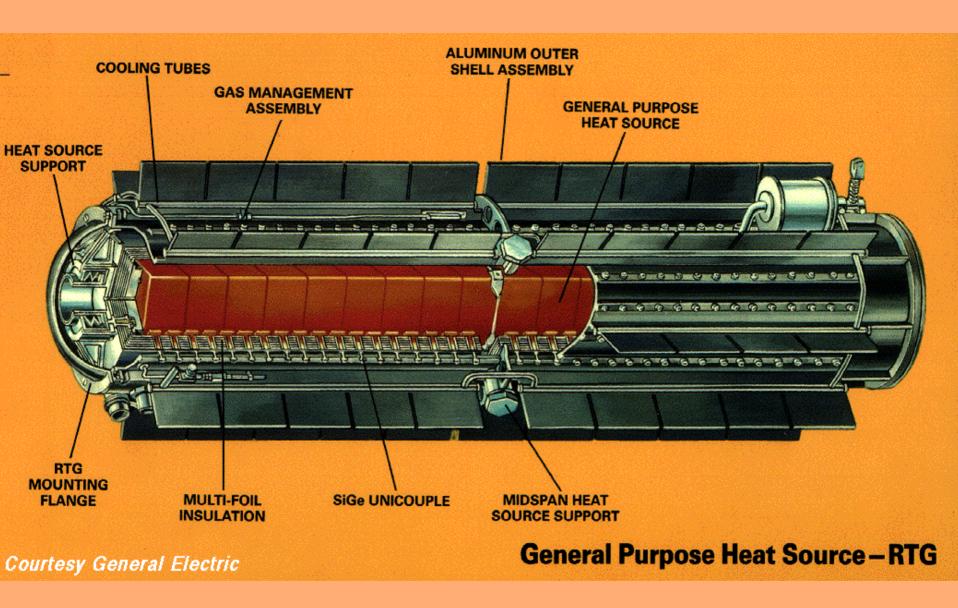
$$\eta_{carnot} = (T_H - T_L)/T_H \approx 50\%$$

$$\eta_{mat} \approx 10-20 \%$$

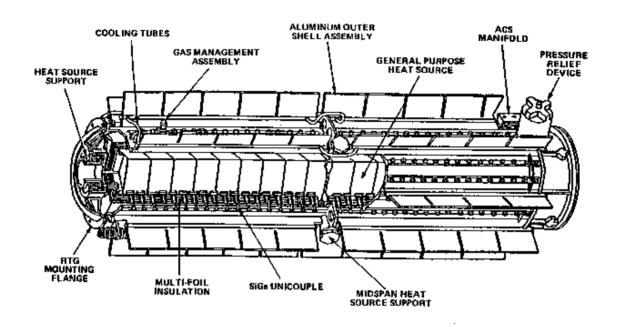
Typical Efficiencies ≈ 5-10%





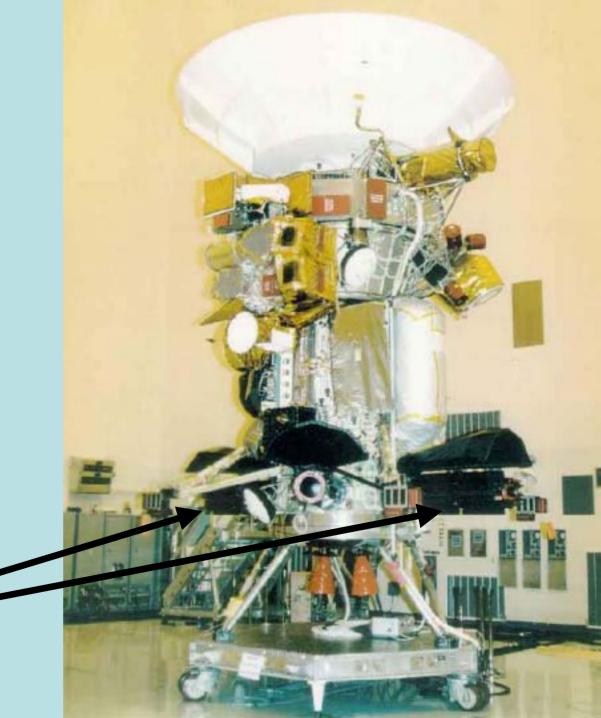


The Galileo RTG Operated Perfectly



- Power Out BOL/EOL = 290/250 W_e
- Mass =55 kg
- Dimensions = 114 cm long/42 cm diam.
- Hot/Cold Junction T °C- 1000/300
- Mass ²³⁸Pu 7.561 kg
- Thermal Power = 4,234 W_t

The Cassini Space Craft

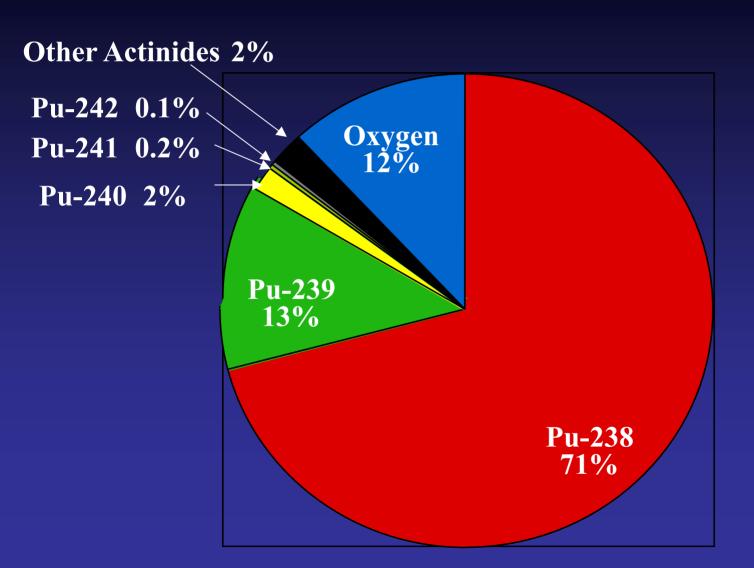


RTG's (3)

Cassini RTG Performance Characterisitics

# of RTG's	3	
Mass/RTG	56 kg (168 kg total)	
Total Power @BOL	888 Watts (electric)	
Total Power @ EOL	628 Watts (electric)	
BOL Thermal Power	13, 182 Watts	
Conversion Efficiency	6.7%	
Mass PuO ₂ /RTG	10.9 kg (32.7 kg total)	
Mass Pu/RTG	9.71 kg (28.8 kg total)	
Mass of ²³⁸ Pu/RTG	7.72 kg (23.2 kg total)	
	(21% of all ²³⁸ Pu already launched)	

Cassini Fuel Composition at Launch

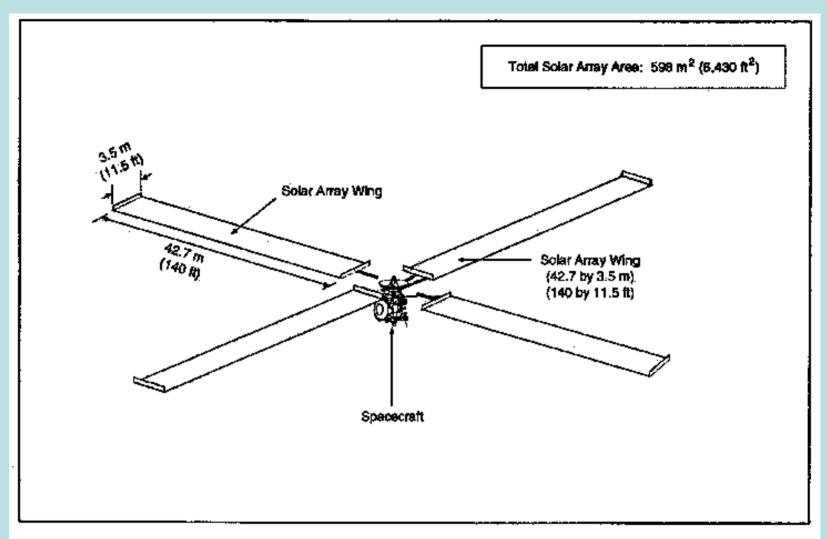


Cassini Electrical Power Requirements

"600-700 W at Saturn (1.6 billion km from sun) for 11 years"

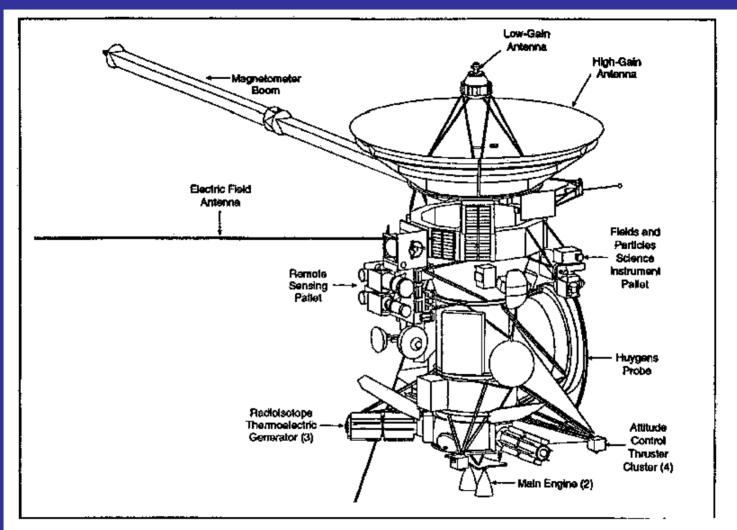
- RTG's
- Mass 168 kg
- Advantages
 - Small size 1.13m x0.43 m dia.
 - No moving parts
 - Easy maneuverability
- Disadvantages
 - Public fear of nuclear

- Solar panels
- Mass 1,337 kg
- Advantages
 - No nuclear material
- Disadvantages
 - No rocket available
 - Slow maneuverability
 - Higher risk of failure



Source: JPL 1994a

FIGURE 2-15. ALL-SOLAR (GRAS APSA) CONFIGURATION FOR THE CASSINI SPACECRAFT



Source: JPL 1993a

FIGURE 2-4. DIAGRAM OF THE CASSINI SPACECRAFT

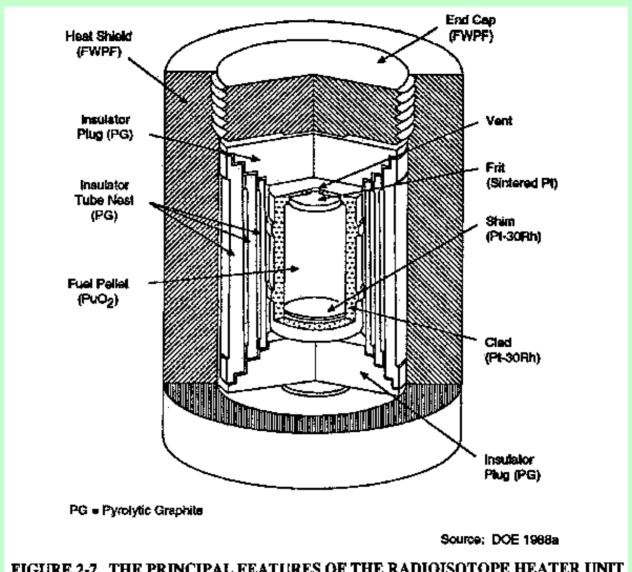


FIGURE 2-7. THE PRINCIPAL FEATURES OF THE RADIOISOTOPE HEATER UNIT

Cassini RHU Performance Characterisitics

# of RHU's	157	
Mass/RHU	40 g (6.28 kg total)	
Thermal Power @BOL	$\approx 1 \text{ Watt}$	
Mass PuO ₂ /RHU	2.7 g (424 g total)	
Mass Pu/RHU	2.38 g (374 g total)	
Mass of ²³⁸ Pu/RHU	1.91 g (300 g total)	

RTG's Have Had a Remarkable Performance Record

# of Launches	# of RTG's	Power /unit, W _e	Mission	Launch Dates
4	4	2.7, 25,25, 30	TRANSIT (navigation)	1961-4, 72
1	2	40	NIBUS (meteorology)	1969
6	6	70	APOLLO(Lunar Exp., 11 ht only)	1969-72
2	8	40	PIONEER-10, 11 (interplanetary)	1972-3
2	4	40	VIKING-1,2 (Mars)	1975
2	4	150	LES (communication)	1976
2	6	150	Voyager-1,2 (Interplanetary)	1977
1	2	275	Galileo (Jupiter)	1989
1	1	275	ULYSSES (Sun)	1990
1	3	296	CASSINI (Saturn)	1997
22	40	4160 tot.		
			Mission failures	
1	1	25	TRANSIT (failed to reach orbit)	1964
1	2	40	NIMBUS (destroyed during launch)	1968
1	1	70	APOLLO-13 (mission aborted)	1970 3
3	4	135 (tot.)		