METEORITES

- LARGLY REPRESENTATIVES OF THE MAIN BELT ASTEROIDS
 BETWEEN MARS AND JUPITER
 - EJECTED BY COLLISIONS COMBINED WITH ORBITAL INTERACTION WITH JUPITER AND SECONDARILY WITH MARS
 LIFE TIMES OF ONLY A FEW MILLION YEARS ONCE IN RESONANCE WITH JUPITER AND MAY DEPLETE SUPPLY TOO FAST

ANISOTROPICALLY EMITTED THERMAL BADIATION (YARKOVFSKY EFFECT) MAY BE ALTERNATIVE MEAN FOR SMALL OBJECT TO AVOID RESONANCE (VOKRUHLICKY AND FARINELLA, 2000, NATURE, 407)

- SOME METEORITES FOR WHICH NO KNOWN ASTEROID
 SPECTURAL TYPE EXISTS
- SOME SPECTURAL TYPES OF ASTEROIDS FOR WHICH NREGOLITH 1M KNOWN METORITES EXIST
 TEMP -14 TO -112 CRATER >10 KM

PHOBOS NASA/JPL 27X22X18KM C-TYPE NREGOLITH 1M TEMP -14 TO -112 CD A TEP > 10 KM

METEORITES AS REPRESENTATIVES OF ASTEROIDS

- STONES: SILICATE DOMINATED (96% OF ALL FALLS) WITH CHONDRITES (88%)
 - PRIMITIPE, UNMELTED, UNDIFFERENTIATED MATERIALS
 - 4.6 B.Y. OLD
 - ABUNDANCES OF ROCK-FORMING ELEMENTS CLOSE TO SOLAR
 USUALLY CONTAIN GLASSY "DROPLETS" CALLED CHONDRULES
- ACHONDRITES (8%)
 - VERY SILICATE-RICH IGNEOUS ROCKS (99% SILICATES AND OXIDES)
 - FORMED BY DENSITY-DEPENDENT DIFFERENTIATION OF SILICATE MAGMA
 - MOST 4.6 B.Y. OLD
 - * STONY-IRONS (1% OF ALL FALLS)
 - ABOUT 50% FERROUS METAL ALLOYS, 50% SILICATES
 - CRYSTALIZED UNDER HIGH PRESSURE
 - IRONS (3% OF ALL FALLS)
 - ABOUT 99% METALLIC FE-NI-CO ALLOYS
 - INCLUSIONS OF FES, PHOSPHIDES, CARBIDES, GRAPHITE, SILICATES, DIAMONDS
 - APPARENTLY RELATED TO HIGH PRESSURE CRYSTALLIZATION

243 IDA NASA/JPL 19X52 KM

METEORITE CHARACTERISTICS

- STONES: SILICATE DOMINATED (96% OF ALL FALLS)
- CHONDRITES (88%)
 - PRIMITIVE, UNMELTED, UNDIFFERENTIATED MATERIALS RELATIVE TO SOLAR ABUNDANCES,
 - 4.6 BY OLD
 - ABUNDANCES OF ROCK-FORMING ELEMENTS CLOSE TO SOLAR
 PROPORTIONS
 - USUALLY CONTAIN GLASSY "DROPLETS" CALLED CHONDRULES
 - ACHONDRITES (8%)
 - VERY SILICATE-RICH IGNEOUS TEXTURED ROCKS (99% SILICATES AND OXIDES)
 - FORMED BY DENSITY-DEPENDENT DIFFERENTIATION, I.E., IN GRAVITY FIELD
 - MOST 4.6 BY OLD

DEIMOS NASA/JPL 15X12X11 KM C-TYPE

METORITE CHARACTERISTICS

STONY-IRONS (1% OF ALL FALLS)

ABOUT 50% FERROUS METAL ALLOYS, 50% SILICATES

• APPARENTLY RELATED TO HIGH PRESSURE CRYSTALLIZATION IN MANTLE OF A NOW DISINTEGRATED PLANET.

RONS (3% OF ALL FALLS)

ABOUT 99% METALLIC FE-NI-CO ALLOYS

 INCLUSIONS OF FES, PHOSPHIDES, CARBIDES, GRAPHITE, DIAMONDS, SILICATES

GEOGRAPHOS NASA/GOLDSTONE ~75X30KM APPARENTLY RELATED TO HIGH PRESSURE CRYSTALLIZATION, SUCH AS IN THE CORE OF A NOW DISINTEGRATED PLANET.

ASTEROIDS IN GENERAL

MAIN BELT ASTEROIDS BETWEEN JUPITER AND MARS

NEAR EARTH ASTEROIDS SOME MAY BE SPENT COMETS

EARTH CROSSING ASTEROIDS SOME MAY BE SPENT COMETS



"CENTAUR" ASTEROIDS BETWEEN JUPITER AND URANUS CHIRON, 1979 VA, AND 133P/ELST-PIZARRO ALSO HAVE COMET-LIKE BEHAVIOR

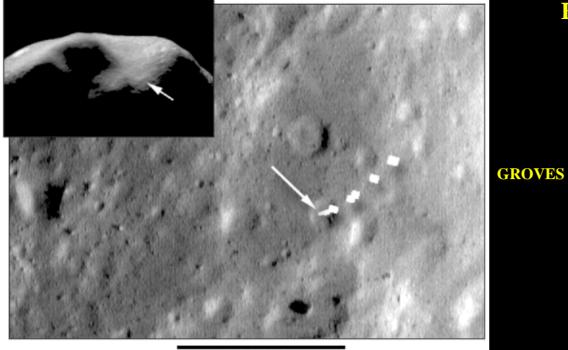
"TROJAN" ASTEROIDS JUPITER'S ORBIT AND CONTROLED BY IT

GENERAL CHARACTERISTICS

ASTEROIDS AND COMETS

RUBBLE PILES (?) NO ASTEROID >150M ROTATES FASTER THAN ONE REVOLUTION PER 2 HOURS CALCULATED LIMIT FOR RUBBLE TO STAY TOGETHER 1998 KY26 IS 30M IN DIAMETER, ROTATES IN 10.7 MIN. AND MAY BE SOLID MAY BE A TRANSITION IN ORBITAL CHARACTERISTICS AND / OR COMPOSITION BETWEEN SOME

> EROS C-TYPE NASA/NEAR SHOEMAKER/APL 11X11X34 KM 1.3 GM/CM³



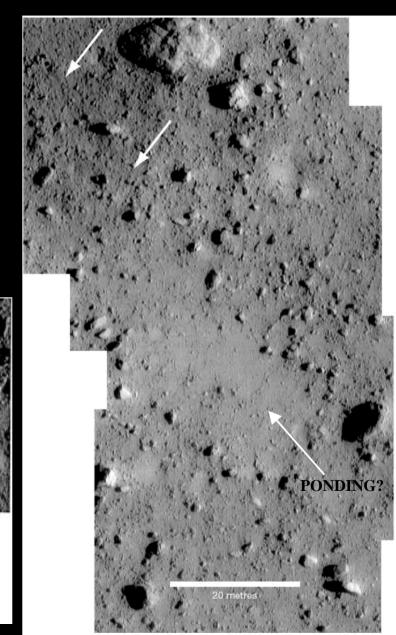
500 metres

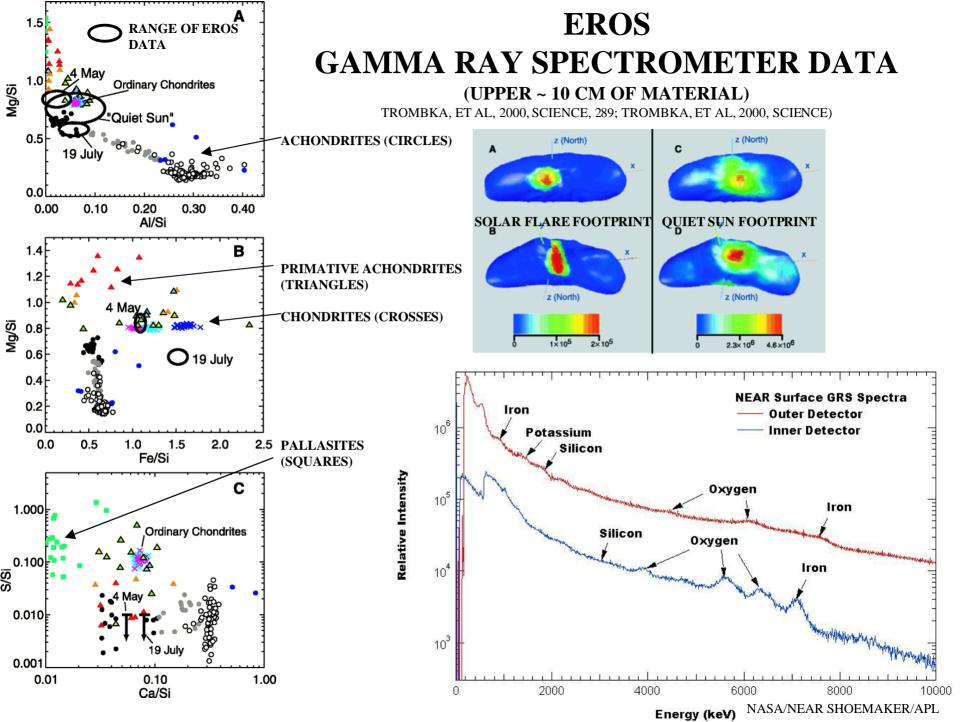
NOTE DEPRESSIONS, I.E., SUSIDENCE FEATURES

5 m 5 m 157417133 15741713 157417198

EROS FINAL DESCENT

NASA/NEAR SHOEMAKER/APL





SPACE WEATHERING

(TENDS TO GIVE A RED TINT TO THE SURFACES OF MOST ASTEROIDS)

WEATHERING FACTORS:

MICROMETEORS (PRODUCE NANO-PHASE IRON)

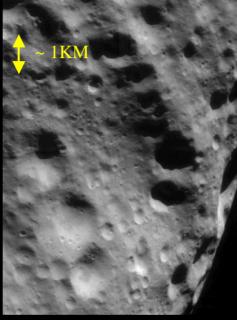
SOLAR WIND/SOLAR FLARE IONS

GALACTIC COSMIC RAYS

COLD / HEAT

EROS NASA/NEAR SHOEMAKER/APL





EROS

SADDLE

CLOSE-UPS AND COLOR

NASA/NEAR SHOEMAKER/APL



CHONDRITES

C-TYPE ASTEROIDS

- 80% OF OBSERVED METEORITE FALLS
- SILICATE-RICH / UNDIFFERENTIATED
 - MONAHANS METEORITE HAS WATER BRINE IN SALT CRYSTAL
- SPECTRA SUGGEST SOURCE MAY BE HEBE IN OUTER MAIN BELT
 - **RIGHT POSITION RELATIVE TO JUPITER**
- 4.567 B.Y. OLD
 - 10⁷ YEAR SPREAD FOR CHONDRULE SOLIDIFICATIO
- **RESEMBLE THE SUN IN COMPOSITIO**
 - EXCEPT IN VOLATILE ELEMENTS
 - OXYGEN ISOTOPES SHOW NON-SOLAR ANOMALIES IN SPINEL, Ca-Al-RICH INCLUSIONS, AND OLIVINE (McKeegan, et al, 1998; Choi, et al, 1998; Hiyagon and Hahimoto, 1999)
- REMNANT MAGNETISM INDICATES FIELD OF 1-10 G
- HIGH PRESSURE SHOCK ASSEMBLAGES IN VEINS

951 GASPRA 19X12/11 KM 7 HR ROTATION PERIOD NASA/GALILEO/JPL

CHONDRITES -2

• CONTAIN "CHONDRULES" RICH IN CA AND AL

ULLIMETER-SCALE IGNEOUS SILICATE

ROUGHLY SPHERICAL, GLASSY, CRYSTALLINE MATERIAL

UP TO 85% of THE MASS OF SOME CHONDRITES

ORIGIN UNCERTAIN
TRANSIENT HEATING EVENTS
POSSIBLY SHOCK HEATING IN THE SOLAR NEBULA BEFORE PLANETESIMALS FORMED
MAY HAVE BEGUN FORMING AT 0.6 AU AND DRIVEN TO 2.5 AU

FIRST STEPS IN TRANSFORMATION OF THE DUST OF THE NEBULA INTO PLANETS (?)

TWO OTHER ASTEROIDS, EUGENIA AND ANTIOPE, AREKNOWN TO HAVE MOONS. 120 KM ANTIOPE CONSISTS OF TWO, EQUAL SIZED BODIES, SEPARATED BY 170KM.

243 IDA (56 KM LONG) AND ITS MOON, DACTYL (1.5 KM) S-TYPE 2.6 GM/CM³

CHONDRITES -3

CHONDRULES CONTAIN "PRE-SOLAR" MATERIAL

(IDENTIFIED BY NON-SOLAR ISOTOPIC RATIOS)

MOST ABUNDANT PRE-SOLAR MATERIAL YET IDENTIFIED

SILICON CARBIDE GRAPHITE NANOMETER-SIZED DIAMONDS REFRACTORY (Al₂O₃) OXIDES SPINEL SILCON NITRIDE METAL CARBIDES

EROS MOSAICS VEVERKA, ET AL, 2000, SCIENCE, 289 NASA/NEAR SHOEMAKER/ARL

5 km

- S-TYPE
 - INNER ASTEROID BELT

OTHER ASTEROIDS

- EVIDENCE OF HEATING AND DIFFERENTIATION
- 29 TELESCOPIC SPECTRA (Binzel, et al., 1996)
 - INTERMEDIATE BETWEEN S-TYPE AND ORDINARY CHONDRITES
 - 1. DISTINCT ROCK TYPES VS DIVERSE LARGER BODIES
 - 2. ABUNDANCE OF OPAQUE MATERIALS
 - **3. FRESH SURFACES (MOST LIKELY)**
- BASALTIC ACHONDRITES (6%)
 - 4 VESTA AT 2.36 AU MAIN BELT PARENT (?)
 - TOUTATIS NEA (RADAR STUDY
 - 4.5X2.4X1.9KM, 2.1 GM/CM3, TWO ROTATIONS, LE., TUMBLING (5.4 AND 7.3 DAYS)
 - 1459 MAGNYA AT 3.15 AU [FRAGMENT OF LARGER BODY (?)]
 - (Lazzaro, et al, 2000, Science, 288)

D-TYPE CARBONACEOUS CHONDRITE (BEYOND MAIN BELT ASTEROIDS

- TAGISH LAKE METEORITE (HIROI, ET AL, 2001, SCIENCE, 293)
 - 4-5% CARBON (MOST KNOWN)
 - PRESOLAR GRAINS
 - CARBONATE MINERALS
- M-TYPE (MAIN BELT)
 - **16 PSYCHE**
 - RADAR SUGGESTS METAL
 - **KLEOPATRA** (Ostro, et al, 2000, Science, 288)
 - RADAR: 217X94X81 KM, DUMBELL SHAPE, 3.5 GM/CM3 REGOMAGM³

EROS C-TYPE (REVISED BY GRS DATA)

11X11X33 KM

5.27 HR ROTATION NASA/NEAR SHOEMAKER/APL

VESTA

BASALTIC A-CHONDRITE (?) MEAN DIA 530KM

460 KM DIAMETER CRATER, 13 KM DEEP MAPPED USING SPECTRAL PROPERTIES

CRATERING ON ASTEROIDS (Veverka, et al, 1997)

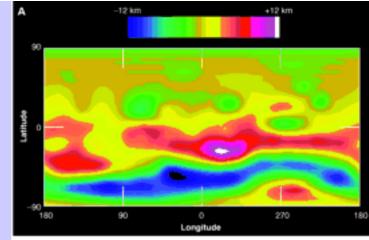
CRATERS FORM WITH DIAMETERS COMPARABLE TO ASTEROIDS MEAN RADIUS

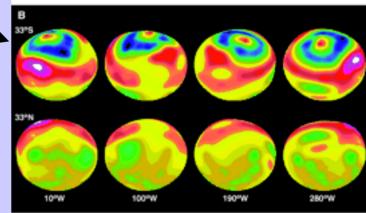
IMPACT DOES NOT BREAK UP BODY AT THIS SIZE

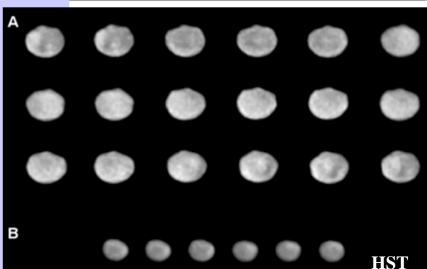
CRATER SIZE-FREQUENCY DISTRIBUTION SIMILAR TO THAT ON THE MOON

LARGE CRATERS HAVE NOT DISTROYED EACH OTHER

PROBABLY DUE TO ACCELERATION OF EJECTA TO ESCAPE VELOCITY







NEAR EARTH ASTEROIDS

- ESTIMATES ARE THAT ABOUT 2000 NEAS EXIST (SEE BOTTKE, ET AL, 2000, SCIENCE, 288)
 - ~950 DETECTED BETWEEN 40 AND 0.01 KM DIAMETER
 - ~900 OTHERS ESTIMATED TO EXIST WITH ~1 KM DIAMETER
 - EJECTED FROM MAIN BELT BY INTERACTIONS WITH JUPITER.
 - COLLISIONS
 - CHAOTIC DYNAMICS INCREASE ORBIT
 - RELATIVELY SHORT (10-100 MYR) LIFE-LIMUS AND THUS MUST BE REPLENISHED RAPIDLY COMPARED TO THE AGE OF THE SOLAR SYSTEM.
 - AMOR TYPE (~29%)
 - ORBIT OUTSIDE THE EARTH'S
 - **APOLLO TYPE (~65%)**
 - ORBIT CROSSES THE EARTH'S.
 - **ATEN TYPE** (~6%)
 - ORBIT INSIDE THE EARTH'S.
- REFLECTANCE SPECTRA INDICATE MANY NEAS ARE SIMILAR TO MAIN BELT ASTEROIDS
 MATH
- OTHERS APPEAR TO BE EXTINCT COMET NUCLEI
 - SURFACE VOLATILES DEPLETED
 - INERT CRUST SEALS REMAINING VOLATILES INSIDE

MATHILDE 59X47 KM ALBEDO 3-4% 17.4DAY ROTATION DENSITY 1.3 C TYPE NASA/NEAR/APL

ECCENTRICITY.

NEAR EARTH ASTEROIDS

- SPECTRA OF NEA 1862 APOLLO
 - METAL, OLIVINE, AND PYROXENE
- 6 TELESCOPIC SPECTRA OF OTHER NEAS
 - SIMILAR TO ORDINARY CHONDRITE METEORITE SPECTA
- ALTERATION IN MANY (HYDROUS, E.G., CLAYS AND IRON OXIDES)
 - BOTH PRE-DATED AND POST-DATED ACCRETION OF PARENT BODY

MATHILDE 59X47 KM C-TYPE ALBEDO 4% (6X<EROS) 1.3 GM/CM³ NASA/NEAR/APL

ASTEROID RESOURCES

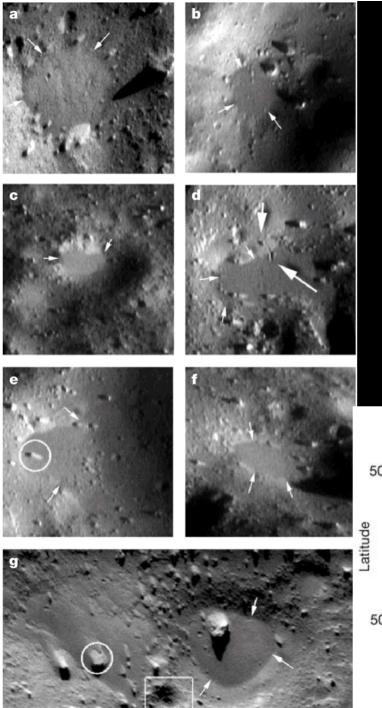
- MAJOR TYPES
 - SILICATE DOMINATED REGOLITH
 - SORTED BY SIZE AND OR DENSITY
 - UNSORTED
 - METAL DOMINATED REGOLITH
 - SILICATE / METAL MIXED REGOLITH
 - SORTED
 - UNSORTED

SILICATE DOMINATED REGOLITH

EROS CLOSE-UP NASA/APL

CHONDRITES (C-TYPE) AND ACHONDRITES

- UNSORTED REGOLITH VERY SIMILAR TO THE MOON'S REGOLITH
 - SOLAR WIND VOLATILES
 - SOLAR WIND DERIVED VOLATILES
 - HYDROUS MINERALS
 - RADIATION PROTECTION MATERIALS
- **EXAMPLES:**
 - EROS [NEAR-EARTH, C-TYPE ASTEROID] (NEAR-SHOEMAKER REFERENCES, E.G., SCIENCE, 2000, 289)
 - LOW DENSITY REGOLITH
 - FINE GRAINED REGOLITH LOCALLY PONDED
 - MATHILDE [NEAR-EARTH, C-TYPE ASTEROID]
 - MAY BE CARBON-RICH [LOW ALBEDO, 1.3 DENSITY]



NEAR SHOEMAKER "PONDED" DEPOSITS NATURE AND DISTRIBUTION

•APPEAR TO BE RESULT OF DOWN SLOPE MOVEMENT

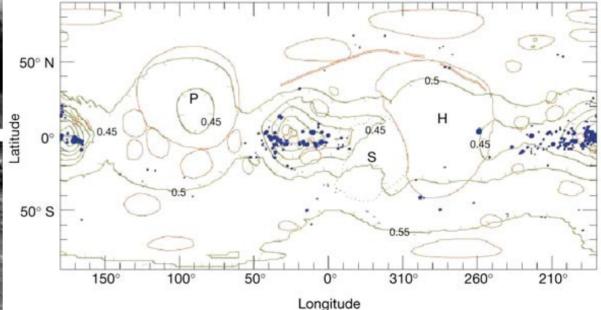
• WHAT ARE THE RESOURCE IMPLICATIONS?

•SIZE DISTRIBUTION?

•DENSITY?

•ELECTROSTATIC PROPERTIES?

NASA/APL: ROBINSON, ET AL, 2001, NATURE, V, 413.



METAL DOMINATED REGOLITH

KLEOPATRA ARECIBO RADAR IMAGESAND RECONSTRUTIONS (OSTRO, ET AL, 2000, SCIENCE, 288

- IRONS (M-TYPE) AND STONY IRONS (S-TYPE)
 - PLATINUM GROUP MUTALS
 - MANUFACTURING METALS
 - SOLAR WIND VOLATILES (?)
- EXAMPLE:
 - KLEOPATRA [MAIN BELT M-TYPE ASTEROID] (Ostro, et al, 2000, Science, 288)
 - RADAR: 217X94X81-KM, DUMBELL SHAPE, 3.5 GM/CM3 REGOLITH
 - POWDERED METAL REGOLITH
 - 1986 DA [NEAR-LARTH M-TYPE ASTEROID]

SILICATE / METAL MIXED REGOLIGH

- STONY IRONS (S-TYPE)
- PROBABLY WOULD COMPLICATE CONCENTRATION PROCESSES
- OTHERWISE, MAY BE BEST FOR SPACE MANUFACTURING
 - DIVERSITY OF PRODUCTS
- EXAMPLE:
 - CASTALIA [EARTH-CROSSING ASTEROID]
 - 2.1 REGOLITH DENSITY

DORMANT COMETS

• HYDROCARBON / DUST CRUST (?)

• ICE-RICH BENEATH CRUST – WATER, HYDROGEN, OXYGEN

ASTEROID RESOURCE ISSUES



- ACCESS TO CAPITAL MARKETS
 - COST OF CAPITAL
 - HIGH RISK = HIGH COST
 - REQUIRES HIGH RETURNS ON INVESTMENT
 - BRIDGE FUNDS TO COVER 10-15 YEAR START-UP WITHOUT A RETURN ON INVESTMENT
 - GOVERNMENT PARTICIPATION (?)
 - EARLY SPINOFF TECHNOLOGY NOT OBVIOUS
- LOW COST LAUNCH ACCESS
 - DEVELOPMENT MIGHT BE SHARED WITH LUNAR ENTERPRISE OR MARS PROGRAM
- RECURRING OPERATIONAL COSTS UNDEFINED
- COST OF 100% RELIABILITY IF AUTOMATED
 - COST OF HUMANS IF NOT AUTOMATED
- OPERATIONAL PROBLEMS
 - VERY LOW GRAVITY
 - ROTATION
- VARIABLE LOCATION OF ASTEROID RELATIVE TO EARTH
- COMPETITION FROM LUNAR RESOURCES
- SIZE OF IN-SPACE MARKET UNCERTAIN
- ECONOMIC IMPACT ON TERRESTRIAL MARKETS FOR PRECIOUS METALS



ASTEROID RESOURCE VALUES

AU

- ASSUME 100 PPM PRECIOUS METAL CONCENTRATION
 - SAME AS SOME METEORITES

Pt

- CURRENT TERRESTRIAL PRODUCTION ~3000 TONNES PER YEAR
 - WORTH ~\$30-40 BILLION PER YEAR
 - NEW SUPPLY THAT COULD UNDERSELL WOULD DEFLATE VALUE
 - SIGNIFICANT WORLD WIDE PRIVATE AND GOVERNMENTAL OPPOSITION TO SUCH COMPETITION FROM SPACE
 - JOBS
 - NATIONAL REVENUE (AUSTRALIA, CANADA, SOUTH AFRICA, RUSSIA, CHILE, ETC.)
- LATER WE WILL COMPARE TO INTRODUCTION OF FUSION POWER BASED ON LUNAR HELIUM-3
 - GRADUAL AND LESS THREATENING ECONOMICALLY
 - FIRST 100KG HELIUM-3 SHIPMENT TODAY WORTH ~\$500 MILLION

ASTEROID RESOURCES SELECTED REFERENCES

- LEWIS, ET AL, 1993, <u>RESOURCES OF NEAR</u> <u>EARTH SPACE</u>, UNIV. ARIZONA PRESS.
- SPACE RESOURCE ROUNDTABLE PUBLICATIONS, M. DUKE, CO. SCHOOL OF MINES.
- INGEBRETSEN, 2001, (REVIEW) IEEE SPECTRUM, AUGUST.
- NEAL, V., ET AL., 1989, (CONSIDERATIONS FOR EVA ON PHOBOS), NASA-17779, SECTION 6.0 http://silver.neep.wisc.edu/~neep602/LEC16/neal.html.
- KARGEL, J.S., 1996, (MARKET VALUES) SPACE 96